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Measurement of the N^{+ 5}S₂ state radiative lifetime

Jiang, Jianzhong, M.S.

University of Nevada, Las Vegas, 1989



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MEASUREMENT OF THE N+ $^5\mathrm{S}_2$ STATE RADIATIVE LIFETIME

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Ву

Jianzhong Jiang

A thesis submitted in partial fulfillment

of the requirements for the degree of

Master of Science

in

Physics

Department of Physics University of Nevada, Las Vegas March, 1989

The thesis of Jianzhong Jiang for the degree of Maser of Science in Physics is approved.

Chairperson, Dr. Victor H.S. Kwong

Examining Committee Member, Dr. James C. Selser

Examining Committee Member, Dr. David P. Shelton

Graduate College Representative, Dr. Stanley Hillyard

Graduate Dean, Dr. Ronald W. Smith

University of Nevada Las Vegas Las Vegas, Nevada June 1989

MEASUREMENT OF THE N^{+ 5}S₂ STATE RADIATIVE LIFETIME

by

Jianzhong Jiang

ABSTRACT

The radiative lifetime of metastable $N^+(2s2p^3 \ {}^5S_2)$ has been measured by directly monitoring the spontanous emission from stored N^+ in a radio-frequency ion trap. The electron excitation technique was used to create $N^+({}^5S_2)$, and the spontaneous emission decay curve via spin-forbidden electric dipole transitions from $2s2p^3 \ {}^5S_2$ to $2s^22p^2 \ {}^3P_{1,2}$ ground state at 2139.68 and 2143.55 Å were observed. A measured value for the 5S_2 radiative lifetime is

 $\tau_{rad} = 6.4 \pm 0.7$ ms

which is in good agreement with the theoretical result by Hibbert and Bates but longer than the measurement by Knight (4.2 \pm 0.6 ms). A careful study of non-radiative quenching and ion loss due to charge transfer of the metastable state was made. It was found that Knight's results were non-reproducible at a higher vacuum system with a meshed wall ion trap. This is perhaps due to unknown impurity gases in Knight's vacuum system. This measurement will clarify the discrepancy between Knight's measurement and theoretical calculations. An accurate value for the lifetime of the metastable N⁺(2s2p³ ⁵S₂) ion is important in interpretation of the UV spectra of aurora.

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I INTRODUCTION

The determination of atomic transition probabilities and lifetimes is an active research area. Various new experimental and theoretical methods have been developed in recent years. New research areas have opened up and the accuracy of the numerical data has significantly improved to meet the increasing demand to understand the basic structure of matter and to create new techniques and devices.¹

Major applications for atomic transition probabilities and atomic lifetimes exist in the following fields:

(1) Astrophysics:

To determine the stellar element abundances, where transition probabilities are the key atomic parameters

(2) Space physics:

Space astronomers need the transition probability data to interpret the far-ultraviolet and soft x-ray spectral line emission from highly ionized species in the solar corona.

(3) Upper atmosphere physics (aeronomy)

Accurate transition probability data for the atmospheric gases are needed for the study of upper atmosphere processes.

(4) Plasma physics, gaseous discharges:

Transition probabilities of stable gases are of interest in the diagnostics of plasmas as well as studies of equilibrium states.

(5) Thermonuclear fusion research

In very hot plasmas, minute heavy element impurities from highly stripped ions radiate large amounts of energy away and thus contribute appreciably to plasma cooling. This is a critical problem in thermonuclear fusion. To analyze and model these energy-loss problems, data for highly stripped ions of wall materials like Cr, Fe, Ni, Mo, and W are needed.

(6) Development of laser systems

Atomic transition probabilities and radiative decay rates of atomic levels are very important parameters needed to assess the potential of a system as a laser, since population inversions can be achieved only if some basic relationships and inequalities are satisfied among these quantities.

This thesis focuses on an experiment that measures $N^+(2s2p^3 \ ^5S_2)$ ion decay via spin-orbit induced electric dipole transitions to the $2s^{2}2p^{2} \ ^3P_{1,2}$ ground state at wavelengths of 2139.68Å and 2143.55Å respectively. This spontaneous decay is believed to be the origin of the λ 2145Å auroral feature.^{2, 3, 4, 5} It is known that auroras are produced by fast ions from the sun that interact with the earth's atmospheric oxygen and nitrogen. Atmospheric molecules such as N₂ are ionized and excited by these fast ions. Those excited ions then decay, radiating light in the characteristic wavelengths of the particular elements.

Considerable interest attaches to the study of the excitation and ionization effects produced in nitrogen by electron impact.^{2, 3, 4, 5, 6, 7, 8} This is partly due to the interpretation of the strong auroral feature near $\lambda 2145$ Å.

Theoretical studies of the $N^+({}^5S_2)$ radiative lifetime were done years ago. A purely theoretical calculation, done by Hibbert and Bates,⁵ yields a lifetime of 6.4 ms. Dalgarno, Victor and Hartquist⁴ extrapolated along the carbon isoelectronic sequence, using previously calculated transition rates for higher ionization states, and obtained a lifetime of 5.8 ms. On the other hand, laboratory study by Knight,³ who measured radiative lifetime of $N^+({}^5S_2)$ ions, gave a lifetime of 4.2 ± 0.6 ms. This value is about 34% off from the theoretical calculations. Measurement by Kwong and Johnson (1982),⁹ and Calamai (1988)¹⁰ consistently gave a higher value of 5.67 ± 0.17 ms and 5.7 ± 0.6 ms respectively. It is hoped that this experiment will resolve the discrepancy among these measurements.

This thesis is divided into several parts. Before presenting experimental details, it will give a theoretical background of many-electron atoms (ions) and a brief description of approximate calculation for the lifetime. Following that the experimental facilities will be introduced and the experiment methods and procedures will be discussed in some detail, including how to estimate some systematic error and nonradiative quenching of the metastable ions. Finally, data will be presented, a final value of the lifetime, including possible systematic errors, will be determined, and a discussion of how the results compare with the former experiment will be given.

II. THEORETICAL OUTLINE

1) The Physics of Many-Electron Atoms (Non-relativistic Theory)

The lifetime of the N⁺(${}^{5}S_{2}$) metastable state is determined by the spin-orbit induced electric dipole transition probability between ${}^{5}S_{2}$ and ${}^{3}P_{1,2}$. The transition rate for such a decay can be described as

$$A_{j} = \overline{A} | \int \Psi_{f_{j}} e \vec{r} \Psi_{i} d\tau |^{2}; \qquad j=1,2$$
(2.1)

where \overline{A} is a constant, Ψ_i , Ψ_{f_j} are the initial and final state functions respectively, and $e \vec{r}$ is the electric dipole operator.

The radiative lifetime of the initial state, then, is simply

$$\tau_{i} = 1 / \sum_{j} A_{j}$$
(2.2)

Therefore, in order to calculate the lifetime of the $N^+({}^5S_2)$, we need to know the related state wave functions. Those wave functions are the solutions of the Schrödinger equation of the system. But the Schrödinger equation cannot be solved exactly for many-electron atoms so that approximation methods must be used. A detailed treatment of a many-electron system should take into account:

- i) H_1 , the kinetic energy of the electrons and their potential energy in the electrostatic (Coulomb) attractive field of the nucleus (assumed to be point-like and infinitely massive compare with the electron mass).
- ii) H_2 , the electrostatic (Coulomb) repulsion between the electrons.
- iii) H₃, the magnetic interaction of the electronic spins with their orbital motion (spin-orbit interaction).
- iv) H_4 , several small effects such as spin-spin interactions between the electrons, various relativistic effects, radiative corrections and nuclear corrections (due to the finite mass of the nucleus, its finite extension, nuclear

magnetic dipole moments, etc.).

In other words, the Hamiltonian of the system can be written as

$$H = H_1 + H_2 + H_3 + H_4 \tag{2.3}$$

If $H_1 \gg H_2 \gg H_3 \gg H_4$, then perturbation theory could be employed. But for many-electron atoms or ions, $H_1 \gg H_2$ may not be true. In that case, the central field approximation should be used before perturbation theory can be employed.

We shall first discuss the central field approximation. For simplicity, let us neglect all the "small" effects: H_3 and H_4 . The Hamiltonian of the N-electron atom (ion) in the absence of external fields can then be written as

$$H = H_{1} + H_{2}$$
$$= \sum_{i=1}^{n} \left(-\frac{h^{2}}{2m} \nabla_{r_{i}}^{2} - \frac{ze^{2}}{(4\pi\varepsilon_{o})r_{i}} \right) + \sum_{i>j=1}^{n} \frac{e^{2}}{(4\pi\varepsilon_{o})r_{ij}}$$
(2.4)

where \vec{r}_i denotes the relative coordinate of the electron i with respect to the nucleus, $r_{ij} = |\vec{r}_i - \vec{r}_j|$ and the last summation is over all pairs of electrons. It is convenient to use atomic units, where $m = h = e = 4\pi\epsilon_0 = 1$. The equation (2.4), then, becomes

$$H = \sum_{i=1}^{n} \left(-\frac{1}{2} \nabla_{r_i}^2 - \frac{z}{r_i} \right) + \sum_{i>j=1}^{n} \frac{1}{r_{ij}}$$
(2.5)

The central field approximation is simply writing the Hamiltonian as follows:

$$H = \sum_{i=1}^{n} \left(-\frac{1}{2} \nabla_{r_{i}}^{2} + V(r_{i}) \right) + \sum_{i>j=1}^{n} \frac{1}{r_{ij}} - \sum_{i=1}^{n} \left(\frac{z}{r_{i}} + V(r_{i}) \right)$$
(2.6)

where

$$V(r_i) = -\frac{z}{r_i} + S(r_i)$$
 (2.7)

and $\sum_{i} S(r_i)$ is a large spherically symmetric component of the inter-electron repulsion term $\sum_{i>j} \frac{1}{r_{ij}}$. From (2.6) we can see that all we have done is to add and subtract the expression $\sum_{i} V(r_i)$ in (2.5). But the perturbation H₂' defined by (2.6) is much smaller than the term $H_2 = \sum_{i>j} \frac{1}{r_{ij}}$ representing the full mutual repulsion between the electrons, therefore perturbation theory can be used.

As a zeroth order approximation, we shall begin by neglecting the perturbation H'_2 and concentrate our attention on the central field Hamiltonian H_c which, as seen from (2.6), contains the kinetic energy $\sum_i -\frac{1}{2} \nabla_{r_i}^2$, the potential energy in the field of the nucleus $\sum_i -\frac{z}{r_i}$, and the average (spherical) electron repulsion energy $\sum_i S(r_i)$. The corresponding Schrödinger equation then reads

$$H_{c}\Psi_{c} = \sum_{i=1}^{n} \left(-\frac{1}{2}\nabla_{r_{i}}^{2} + V(r_{i}) \right) \Psi_{c} = E_{c}\Psi_{c}$$
(2.8)

where Ψ_c is the N-electron central field wave function which satisfies the requirements of the Pauli exclusion principle and can be written as

$$\Psi_{c}(q_{1},q_{2},\cdots q_{n}) = \frac{1}{\sqrt{n!}} \begin{vmatrix} \mu_{\alpha}^{(q_{1})} & \mu_{\beta}^{(q_{1})} & \dots & \mu_{\gamma}^{(q_{1})} \\ \mu_{\alpha}^{(q_{2})} & \mu_{\beta}^{(q_{2})} & \dots & \mu_{\gamma}^{(q_{2})} \\ \vdots \\ \mu_{\alpha}^{(q_{n})} & \mu_{\beta}^{(q_{n})} & \dots & \mu_{\gamma}^{(q_{n})} \end{vmatrix}$$
(2.9)

where q represents the spatial and spin coordinate and

$$\mu_{v}(q_{i}) = \mu_{nlm_{1}m_{s}}(q_{i}) = \mu_{nlm_{l}}(\vec{r}_{i}) \chi_{m_{s}}$$
(2.10)

Since H_c is central and separable to N single-independent Hamitonian, so the

one-electron or central field orbitals $\mu_{v}(q_{i})$ are products of spin orbitals times spatial orbitals, which are products of radial functions times spherical harmonics,

$$\mu_{n1m_{1}}(\vec{r}_{i}) = R_{n1}(r_{i}) Y_{1m_{1}}(\theta, \phi)$$
(2.11)

where the radial function satisfies the equation

$$-\frac{1}{2}\left(\frac{d^2}{dr_i^2} + \frac{2}{r_i}\frac{d}{dr_i} - \frac{l(l+1)}{r_i^2}\right)R_{nl}(r_i) + V(r_i)R(r_i) = E_{nl nl}(r_i)$$
(2.12)

and $Y_{lm_1}(\theta, \phi) \chi_m$ for every electron is known. The total energy E_c in the central field approximation is the sum of the individual electron energies, namely

$$E_{c} = \sum_{i=1}^{n} E_{n_{i}l_{i}}$$
(2.13)

in which, the energy eigenvalues of each electron, $E_{n_i l_i}$, do not depend on the quantum number m_l , because the potential $V(r_i)$ in (2.7) is spherically symmetric. However, in contrast to the hydrogenic case, they depend on both n and l because $V(r_i)$ is not simply the Coulomb potential. Higher order approximation can be made through perturbation theory.

A more elaborate approximation for complex atoms (ions) is the Hartree–Fock or self–consistent field method.³⁵ This is an independent particle model, according to which each electron moves in an effective potential which takes into account the attraction of the nucleus and the average effect of the repulsive interactions due to the other electrons. Each electron in a multielectron system is then described by its own wave function. In accordance with the independent particle approximation and the Pauli exclusion principle, the N–electron wave function is a Slater determinant Ψ (same as Eq.(2.9)), or, in other words, an antisymmetric product of individual electron spin–orbitals. The Slater determinant is optimized by using the variational method to determine the 'best' individual electron spin–orbitals. 2) Calculations of the $N^+({}^5S_2)$ Radiative Lifetime

The ${}^{5}S_{2}$ state is the first excited state of N⁺(1s²2s 2p³) above the terms of the (1s²)2s²2p² configuration: ${}^{3}P$, ${}^{1}D$, ${}^{1}S$. Direct radiative decay from ${}^{5}S_{2}$ to those lower energy states can only occur through magnetic quadrupole transitions. However, the decay could also be possible through spin-forbidden electric dipole transitions (intersystem transitions). We can achieve this transition if the initial ${}^{5}S_{2}$ state is slightly mixed with other J = 2 levels (${}^{3}D_{2}$, ${}^{3}P_{2}$, ${}^{1}D_{2}$, ${}^{1}P_{2}$) of the 1s²2s 2p³ configuration through spin-orbit interaction.

Hibbert and Bates⁵ proved that the spin-orbit induced electric dipole transitions from ${}^{5}S_{2}$ to ${}^{3}P_{1}$ and ${}^{3}P_{2}$ dominated all others, and that the principal magnetic quadrupole transitions, also to the ${}^{3}P_{J}$ levels, had probabilities approximately five orders of magnitude smaller. Based on the Hartree-Fock functions of the $2s^{2}2p^{2}$ and $2s 2p^{3}$ states made by Roothaan and Kelly,¹¹ and then by introducing further configurations Hibbert and Bates determined that the lifetime for the ${}^{5}S_{2}$ is 6.4 ms.

Dalgarno et. al,⁴ on the other hand, estimate a lifetime of 5.8 ms by extrapolation along the isoelectronic sequence. Our experimental result $(6.4 \pm 0.7 \text{ms})$ is in very good agreement with these results.

III. EXPERIMENTAL FACILITIES

A block diagram for the facility to measure the $N^+({}^5S_2)$ lifetime is shown in Fig.(1). The facility includes a vacuum system, a rf-quadrupole ion trap, an ion creation and detection system, a photon detection and counting system, a timing control system, and a data analysis system.

1) The Vacuum System

The vacuum system produces a background pressure low enough to minimize the influence of impurity ion signals on our experimental results.

As can be seen in Fig.(2), the vacuum chamber consists simply of a main chamber to house the ion trap apparatus, electron gun and CEM, and a flange to which the ion trap apparatus is attached. The entire chamber is pumped by a VHS-4 diffusion pump backed by a mechanical pump. A liquid nitrogen cold trap is placed between the chamber and the diffusion pump to minimize contamination of the vacuum chamber by pump oil. A gate valve with a thermocouple gauge is set in between the diffusion pump and mechanical pump to monitor the background pressure of the system before the diffusion pump can be started.

The vacuum chamber configuration is shown in Fig.(3). The main chamber consists of six ports. The bottom port is connected to the vacuum pump system. The top port is simply sealed (it is not used in this experiment). The right port is connected to the photomultiplier tube with a sapphire window and a filter in between. Two side ports are used for mounting the ion trap apparatus with a feedthrough (on the lower end) and the channel electron multiplier (on the higher end). The alignment between the ion trap and PMT is important for correct photon detection.



Fig.(1) Block diagram of the $N^+({}^5S_2)$ lifetime experiment.

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With this vacuum system, a background pressure of $5 \times 10^{-9} \tau \text{orr}$ can be obtained without baking. The pressure in the vacuum chamber can be measured by a quadrupole mass spectrometer (Masstorr DX100)³⁶ mounted on the left port of the vacuum chamber.

An interlock protection system is used which will turn off the diffusion pump and isolate the vacuum chamber to avoid pump oil from the mechanical pump if:

i) the power supply is shut off for more than 15 sec.

ii) the cooling water for diffusion pump runs below an acceptable level.

iii) unacceptable high pressure (10 mTorr) at the mechanical pump is reached.

2) The RF-Quadruple Ion Trap

The ion trap is the key device for this experiment. It has the ability to confine the desired ions for relatively long time with negligible perturbation during the measurement. The ion trap in this work was made from meshed stainless steel. The configuration of the ion trap used in this work is shown in Fig.(4). The trap is composed of two end electrodes and a ring electrode with small holes cut into the ring electrode to allow better photon detection. The size of this ion trap (see Fig.(5)) is

$r_{o} = z_{o} = 1.68 \text{ cm}$

The quadrupole ion trap was first described by Paul¹² and its feasibility was first demonstrated by Berkling¹³ and Fischer.¹⁴ Since that time, many workers have contributed to the development of the quadrupole trap and the study of its properties and characteristics. Most notable among these are Dawson <u>et al.^{15, 16, 17, 18, 19, 20</u>, ^{21, 22} and Todd <u>et al.^{19, 23, 24, 25</sub> who carried out extensive investigations on many aspects of the trap. However, the studies mentioned above concentrated on ion traps with hyperbolic electrodes.</u>}</u>}

In 1973, Benilan and Audoin²⁶ presented the theoretical description of a



Fig.(4) Configuration of the Ion Trap.

cylindrical ion trap. The cylindrical ion trap was chosen not only because it is easy to machine but because it was felt that the various holes that had to be cut into the trap electrodes would perturb the electric fields to a smaller degree in the cylindrical trap than in the hyperbolic trap.

To understand the performance of the ion trap, a brief outline of ion trap theory is necessary. Detailed discussion of ion traps can be found in many other papers.^{27, 28} The trap in this work, as mentioned above, is a cylindrical ion trap. Unlike a hyperbolic trap, there is no simple solution for the potential distribution. However, it has been shown that the potential distribution at and near the center of the cylindrical trap is very close to that of the hyperbolic trap.²⁶ Furthermore, Knight²⁹ already gave the general form of the potential distribution, in which the axial size z_0 is independent to the radial size r_0 , and it was well confirmed by our experiments as well as other experiments. Since the ions are confined near the center of the trap, so, for simplicity, the theoretical discription for hyperbolic trap can be used for cylindrical trap. The potential distribution, then, is

$$\phi(\mathbf{r}, \mathbf{z}) = \frac{U}{r_0^2 + 2z_0^2} \left(\mathbf{r}^2 - 2(\mathbf{z}^2 - \mathbf{z}_0^2) \right)$$
(3.1)

where we specified that the two end electrodes are grounded and there is a potential

$$U = U_{o} - V_{o} \cos\Omega t \tag{3.2}$$

on the ring electrode as shown in Fig.(5). When a charged ion is in this potential it will see a electric field given by

$$\mathbf{E} = -\nabla \phi(\mathbf{r}, \mathbf{z}) \tag{3.3}$$

and a force

$$\vec{\mathbf{F}} = \mathbf{e} \, \vec{\mathbf{E}} = \mathbf{m} \, \vec{\mathbf{r}} \tag{3.4}$$

acting on it. So the equations of motion for a mass m with charge e become:



• • • •



$$\ddot{r} = -\left(\frac{e}{m}\frac{2U_{o}}{r_{o}^{2} + 2z_{o}^{2}}\right)r + \left(\frac{e}{m}\frac{2V_{o}}{r_{o}^{2} + 2z_{o}^{2}}\right)r\cos\Omega t$$
(3.5)

$$\ddot{z} = \left(\frac{e}{m}\frac{4U_{o}}{r_{o}^{2} + 2z_{o}^{2}}\right)z - \left(\frac{e}{m}\frac{4V_{o}}{r_{o}^{2} + 2z_{o}^{2}}\right)z\cos\Omega t$$
(3.6)

If we define:

$$2a_{r} = -a_{z} = \frac{16eU_{o}}{m(r_{o}^{2} + 2z_{o}^{2})\Omega^{2}}$$
(3.7)

$$2q_{r} = -q_{z} = \frac{8eV_{o}}{m(r_{o}^{2} + 2z_{o}^{2})\Omega^{2}}$$
(3.8)

and

$$v = \frac{\Omega t}{2} \tag{3.9}$$

then both r and z equations of motion are of the form

$$\frac{d^2 u}{dv^2} + (a - 2q\cos 2v) u = 0$$
(3.10)

This is the Mathieu equation. A class of solution, depending on two parameters a and q, leads to finite amplitude of the ion motion, i.e., to confinement. Fig.(6) shows the stability diagram obtained by requiring the values of a_z , a_r , q_z , and q_r to be such that the motion is stable in both the axial and radial directions. For an ion of fixed $\frac{e}{m}$, selection of the trap voltages and frequency (U_o, V_o, and Ω) can locate the operating point on this diagram. Only if the operating point specified by a and q is in the interior of the closed region of Fig.(6) can trapping occur. An approximate solution to the equation of motion, valid for $q \le 0.4$ and $a \ll q$ can be obtained by dividing the motion into two components,³⁰ a displacement, δ , due to the micromotion resulting from the high frequency field and a larger displacement, \bar{u} , which describes the extent of the motion averaged over a period of the rf drive potential, namely:

$$\mathbf{u} = \overline{\mathbf{u}} + \mathbf{\delta} \tag{3.11}$$

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Fig.(6) Stability diagram of the ion trap

where $\delta \ll \overline{u}$ but $\frac{d\delta}{dt} \gg \frac{d\overline{u}}{dt}$.

Then Mathieu's equation becomes

$$\frac{d^2}{dv^2} (\bar{u} + \delta) + (a - 2q\cos 2v) \ \bar{u} = 0$$
 (3.12)

or

$$\frac{d^2\delta}{dv^2} = -(a - 2q\cos 2v)\,\overline{u} \tag{3.13}$$

Since $a \ll q$ and \overline{u} changes slowly, integration of the above equation gives

$$\delta = -\frac{q\bar{u}}{2}\cos^2\nu \qquad (3.14)$$

and

$$\mathbf{u} = \overline{\mathbf{u}} + \delta = \overline{\mathbf{u}} - \frac{q\overline{\mathbf{u}}}{2}\cos 2\mathbf{v} \tag{3.15}$$

Substitute this into the original Mathieu equation, we have

$$\frac{d^2 u}{dv^2} = -a\bar{u} + \frac{aq\bar{u}}{2}\cos^2\nu + 2q\bar{u}\cos^2\nu - q^2\bar{u}\cos^22\nu$$
(3.16)

Averaging over one cycle of the drive frequency gives the equation for the slow macromotion (since the fast motion due to rf drive averaged over a period of the rf drive is equal to zero.)

$$\frac{d^2 \bar{u}}{dv^2} + (a + \frac{q^2}{2}) \,\bar{u} = 0 \tag{3.17}$$

which, written in terms of time, becomes

$$\frac{d^2 \overline{u}}{dt^2} + (a + \frac{q^2}{2}) \frac{\Omega^2}{4} \overline{u} = 0$$
(3.18)

This equation corresponds to simple harmonic motion with the oscillation frequency

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$$\omega = \frac{\Omega}{2} \left| a + \frac{q^2}{2} \right| < \Omega$$
 (3.19)

$$\overline{u} = A\cos\omega t$$
 (3.20)

and

$$u = \overline{u} + \delta = A\cos\omega t \left(1 - \frac{q}{2} \cos\Omega t \right)$$
(3.21)

A reasonable way to picture the ion motion, then, is a slow harmonic motion at frequency ω upon which is superimposed a micromotion at the drive frequency Ω . Note that the spectrum of the motion contains components at frequency ω as well as $\Omega \pm \omega$, which arise from the cos ω t cos Ω t term.

The force (averaged over one cycle of drive frequency) on an ion of mass m and charge e is therefore given by

$$m\frac{d^2\overline{u}}{dt^2} = -m\left(a + \frac{q^2}{2}\right)\frac{\Omega^2}{4}\overline{u} = -e\frac{d\overline{D}}{du}$$
(3.22)

For the motion along z direction:

$$\frac{d\overline{D}_z}{dz} = \frac{m}{e} (a_z + \frac{q_z^2}{2}) \frac{\Omega^2}{4} z$$

$$= \left(\frac{-4U_{o}}{(r_{o}^{2} + 2z_{o}^{2})} + \frac{8eV_{o}^{2}}{m(r_{o}^{2} + 2z_{o}^{2})\Omega^{2}}\right) z$$
(3.23)

$$D_{z} = \left(\frac{4eV_{o}^{2}}{m(r_{o}^{2} + 2z_{o}^{2})^{2}\Omega^{2}} - \frac{2U_{o}}{(r_{o}^{2} + 2z_{o}^{2})}\right) z_{o}^{2}$$
(3.24)

similarly, for the r motion we have

$$D_{r} = \left(\frac{eV_{o}^{2}}{m(r_{o}^{2} + 2z_{o}^{2})^{2}\Omega^{2}} + \frac{U_{o}}{(r_{o}^{2} + 2z_{o}^{2})}\right)r_{o}^{2}$$
(3.25)

where D_r and D_z are defined as potential well depths in the radial and axial directions.

The total effective potential seen by an ion, then, can be written as

$$\phi_{eff} = D_r \frac{r^2}{r_0^2} + D_z \frac{z^2}{z_0^2}$$
(3.26)

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to make $\phi_{eff} = 0$ at the trap center.

In the case of a spherical well, we have $D_r = D_z = D$. For our trap design $r_0 = z_0$. We can get

$$D = \frac{2}{3} U_{o}$$
 (3.27)

which shows that the spherical potential well depth is directly proportional to the DC bias of the ion trap.

The optimum trapping parameters have been investigated both theoretically and experimentally.^{31, 32} The results, plotted on the stability diagram in Fig.(7), are in good agreement with each other. They are also in good agreement with the data obtained by our trap. For a fixed q_z (or a_z), the value of a_z (or q_z) which results in a spherical potential well will give maximum stability.

3) Electron Gun and Channel Electron Multiplier

The ion creation and detection system in this experiment includes an electron gun and a channel electron multiplier.

A. Electron Gun

The electron gun used was a 7mm long, 3mm diameter tungsten dispenser cathode, which was made from porous tungsten with a formula of barium oxide dispersed throughout the matrix. This dispenser cathode will allow higher emission current densities at lower operating temperature with decreased sublimation rates and greater resistance to poisoning from the cathode environment.

The electron gun was screw fixed to a stainless steel support and mounted 2mm from the lower end cap of the ion trap [Fig.(3)]. The cathode was heated by $I \simeq 0.8A$, 60Hz current with the potential difference of 4.0V across it during the experiment. In order to generate $N^+({}^{5}S_2)$ from background N₂ gas through electron impact with molecular nitrogen, bias voltage of -200V was set to the electron gun whereas the lower end cap of the ion trap was grounded. Under these circumstances



Fig.(7) Contours of the number of ions stored

(from Reference 28)

the electrons produced by the electron gun are accelerated into the ion trap and bombard the molecular nitrogen there. Alternatively, the electrons can be confined to the region around the electron gun by adding a positive bias of +100v to the electron gun. This prevents the unwanted electrons getting into the ion trap to disturb the lifetime measurement.

B. Channel Electron Multiplier (CEM)

The model 4716 high current CEM has been used to detect the ions confined in the ion trap. The observation takes place during the electron gun "off" time. A negative high voltage pulse (-70v) was applied to the higher end cap electrode of the trap nearest to the CEM so that the ions confined in the ion trap were driven out through this electrode, and into the CEM, as seen in Fig.(3). The ion signal was amplified by means of secondary electron emission in the CEM, and then displayed on an oscilloscope. We confirmed that the signals could be ascribed to N⁺ or N₂⁺ since they vanished if either the N₂ was removed or the ion trap was detuned to prevent their storage. Time of flight technique has also been used to confirm N⁺ storage.

4) The Photon Detection System

The photon detection system consists of i) a sapphire window to allow UV photons to pass through. ii) an interference filter, which transmits 21.6% at a wavelength of 2150 Å with a bandwidth of 230 Å, to attenuate blackbody radiation from the hot electron gun cathode. iii) a Photomultiplier Tube (PMT) to detect photons, and iv) a Multi-Channel-Analyzer (MCA), which serves as a photon counter. These are shown schematically in Fig.(8).

The PMT converts the incident radiation into electrical signals by use of the phenomenon of photoemission and then amplifies the signals by means of secondary emission.



Fig.(8) Photon Detection System

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The PMT used in this work is the EMR Model 541Q-05M-13 multiplier phototube, which is an end-on tube with a 25mm effective diameter, semitransparent, solar blind photocathode. Its window material is selected ultraviolet grade sapphire (Al₂ O₃). The spectral sensitivity is characterized by high ultraviolet sensitivity with an extremely sharp cut off of atmosphere-penetrating solar radiation. As shown in Fig.(9), the short wavelength cut off is at 1700Å and the longer wavelength response is extended out to 3650Å with quantum efficiency of 0.001%. Peak sensitivity occurs near 2537Å. Typical voltages of 2100v to 2600v are required for current amplifications of 10^6 to 10^7 .

The output pulses from the PMT are amplified by a 9301 ORTEC Fast Amplifier., passed through a discriminator (ORTEC 4890 Pre–Amp SCA) and routed to a scaler (MCA) which is gated on by the falling edge of the CPU relaxation pulse. The pulse amplification gain is set to give saturating pulses at a moderate high voltage and the discriminator is set higher than the amplifier noise level. The MCA performs in MCS (Multichannel Scaling) mode, which is a time sweep of the channels in the MCA with each channel being an equal time interval to total sweep. During each channel time interval, the memory location for that channel counts digital pulses at random rates up to 10 MHz. Thus, the resulting display is a frequency histogram with time as the horizontal axis. The dwell time ranges from 10 μ s through 1 second.

The data received by the MCA can either be displayed on a built-in Cathode-Ray Tube (CRT) or stored in the non-volative memory. The stored data can be recalled when desired.

5) Control Process Unit (CPU)-the Experimental Timing Control Box

In the previous sections, the various functions and measurements which make up the experiment were discussed separately. An understanding of how these pieces



Fig.(9) The Photomultiplier Tube spectral response.

are integrated to form the whole experiment is best accomplished by describing the experimental timing control box (CPU).

The CPU contains logic circuitry which generates 4 TTL logic pulses. The first pulse controls the electron gun "on" time (the ionization period). The trailing edge of the first pulse triggers the second pulse, the relaxation pulse. The length of the relaxation pulse determines the waiting period between electron gun shut off and the detection pulse. The third pulse, the detection pulse, is turned on by the falling edge of the relaxation pulse. The leading edge of this pulse turns on the CEM. The measurement is allowed after an appropriate relaxation time and before the next electron gun "on" pulse. The fourth pulse, triggered by the trailing edge of the detection pulse, will dump the ions out from ion trap to the CEM. The dump pulse and the CEM are used only to identify the ions confined in the ion trap. This will help to select appropriate parameters for the ion trap to confine the desired ion. During the $N^{+}(^{5}S_{2})$ radiative lifetime measurement, the MCA is gated on by the leading edge of the detection pulse to gather the data, whereas the CEM and the dump pulse are not used to avoid the disturbance of ion storage. The MCA sweep time duration has to be longer than the CPU detection pulse width, otherwise the MCA sweep signal will turn on automatically and start to sweep again if the CPU detection pulse is still high. On the other hand, the MCA sweep should turn off before the next e gun pulse for the same reason.

6) Data Analysis System

Data analysis was done by using the "DISCRETE" program.³³ It is a FORTRAN IV program for the automatic analysis of data being represented by:

$$N_{k} = \sum_{j=1}^{m} \alpha_{j} \exp(-\lambda_{j} t_{k}), \quad k = 1, 2, \cdots N, m \le 9$$

A provision can be made for an unknown baseline component α_o with

 $\lambda_o = 0$. It is completely automatic in that only the raw data(i.e., the N_k and t_k) are input; no potentially biased initial guesses at the α_j , λ_j or m are needed or even allowed. So the results from this data analysis should be more reliable.

IV EXPERIMENTAL METHODS AND PROCEDURES

1) Experimental Method and General Procedures

The method used to measure the $N^{+5}S_2$ state lifetime was simply to detect spontaneously emitted photons as a function of time.

A single exponential decay signal can be characterized by

$$N(t) = N_o \exp(-t/\tau)$$
(4.1)

and the decay rate at any instant of time t is defined by

$$\frac{1}{\tau} = -\frac{d}{dt} \left(\ln N(t) \right)$$
(4.2)

substituting (4.1) into (4.2), we have

$$\ln N(t) = \ln N_o - \frac{1}{\tau} t$$
(4.3)

So the graph of $\ln N(t)$ vs t is a straight line with slope $-\frac{1}{\tau}$ independent of t, and τ is the radiative lifetime. Experimentally we have two steps to complete before the lifetime measurement. First of all, we have to create enough N⁺(⁵S₂) ions, but when we create N⁺ we create other kinds of ions as well. So the second step is to select those desired ions and store those ions in a disturbance free area for a time which is much longer then the expected lifetime. Here disturbance free means the disturbance from the surroundings (e.g. the ion trap's RF field) are negligible compared with the interaction between nucleus and electrons of the ion. The third step, then, is simply to detect the spontaneous emission while the ions are in storage. The final step is data analysis. In this experiment, the first three steps were controlled by CPU TTL signals with the typical timing sequence performed by CPU as shown in Fig.(10). The data analysis was performed later by "DISCRETE" computer program. The procedures, described in some detail, are following.

2) Creation of N⁺

N⁺ ions, in both metastable and ground states, were created in the ion trap



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due to the electron impact dissociation and ionization of N_2 gas. The process can be described as

$$N_2 + e \rightarrow N^+ + N + 2e \tag{4.4}$$

where the effective threshold to produce $N^+({}^5S_2)$ is $37eV.^4$ The kinetic energy of the electrons, before impact with N₂, is determined by the negative bias between the e gun and the end cap of ion trap plus the rf potential field in the ion trap. This initial electron kinetic energy can determine the rates for creation of $N^+({}^5S_2)$ from a certain N₂ back pressure. We estimated those rates by monitoring the 2140Å photon signal while we scanned the negative bias and finally set the bias at -200V. The electron energies under the trap RF power of $V_o = 400V$, $f_o = 1.2MHz$, then, can be estimated to be under 500eV. In this energy range, the cross section of producing $N^+({}^5S_2)$ via electron impact on N₂ has been estimated as $\ge 1 \times 10^{-18}cm^2$. ^{3, 4} Time of flight measurements of ion velocities following dissociative ionization of N₂ suggest that $N^+({}^5S_2)$ is formed with an initial kinetic energy around $3eV.^{8, 34}$

The electron gun negative and positive bias, as mentioned in section III, were supplied by a dual regulated power supply. The switching time (from -200V to +100V) was controlled by the CPU e gun signal. The e gun "on" time (for which -200V was set between electron gun and the end cap of ion trap) was long enough to produce a number of N⁺(⁵S₂) ions in the ion trap. To determine the number of N⁺(⁵S₂) ions produced in time T, the following equation is used:

$$N = I\sigma n l T \tag{4.5}$$

where I is the electron beam current, n is the neutral parent gas number density, σ is the ionization cross section, 1 is the path length of the electrons, and T is the electron gun "on" time. In this experiment:

$$I \sim 5A = 5 \times 10^5 \times 6.2 \times 10^{13} \text{e/sec} = 3 \times 10^{19} \text{e/sec}$$

as tested by Kwong (1981)⁹

n ~ 10⁻⁶Torr ×
$$\frac{2.69 \times 10^{19}}{760}$$
 = 3.54 × 10¹⁰ /cm³
 σ ~ 10⁻¹⁸cm²
l ~ 2 cm

and in this experiment we choose $T = 0.5ms = 5 \times 10^{-4}sec$. The number of ions produced, then, is

This number is much greater than the number of ions (~ 10^5) that could be confined in our ion trap due to space charge effects.

The CPU electron gun signal and relaxation signals, combined with the dual regulated power supply's response signal are shown in Fig.(11). From there we can see that the power supply's response signal has a significant rise time ($\simeq 1.2$ ms). Although the electron gun control pulse has already turned off, some energetic electrons still can get into the ion trap to create some "noise" during this period of time. Increasing the positive bias can decrease the rise time. The relaxation signal has been extended such that the relaxation time is greater than the rise time + 0.3ms (which is the fast decay time attributed to emission in the N₂ Lyman–Birge–Hopfield bands³).

In order to avoid unexpected disturbances, the vacuum was kept very high $(0.5 \times 10^{-8}\text{Torr})$ before filling the N₂ gas. The composition of residual gas at background pressure $5 \times 10^{-9}\text{Torr}$ in the vacuum chamber is shown in Fig.(12). It includes H, H₂, HO, H₂O, CO, CO₂, etc. A background run was made without fill N₂, whereas the other parameters were all set the same as those for the N⁺(⁵S₂) lifetime measurement. The PMT did not see any decay under this circumstance. That means that the radiation from the residual gas is negligible. Ultrahigh purity N₂ gas (99.9999%) was used as a parent gas for creation of N⁺.



Fig.(11) The CPU e gun signal, relaxation signal combined with the dual regulated power supply's response signal

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Fig.(12) The residual gas composition

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3) Ion Selection and Storage

Ions created via electron impact with N_2 could be in many different states. It is very important to exclude the ions other than N^+ from storage and to confine only N^+ ions in a perturbation-free environment for times much longer than the expected radiative lifetime of a few milliseconds. In the last chapter, we saw that the rf quadrupole trap is an ideal tool to do this job.

Calculations locating various ion density peaks on the stability diagram were made. The results showed that:

- i) at $V_o = 400V$, $f_o = 1.2MHz$, $U_o = 25V$; N⁺ ions sit right on the spherical potential well line (see Fig.(13 a)) and at the largest density area, where $q_z \sim 0.5$, $a_z \sim -0.04$. This will give the maximum stability and largest ion density for N⁺ confinement. On the other hand, N₂⁺ and N⁺² can not be confined under this condition.
- ii) at $V_o = 255V$, $f_o = .65MHz$, $U_o = 20V$; only N_2^+ can be trapped. (see Fig.(14 a))
- iii) at $V_o = 220V$, $f_o = .65MHz$, $U_o = 20V$; both N⁺ and N⁺₂ can be confined. (see Fig.(15 a))

These results are for a trap with ideal hyperbolic electrodes. However, the potential surfaces of a cylindrical ion trap near the center of the trap are very close to the ideal one, as mentioned in the last chapter. To test the system, the ion trap voltages and frequencies have been set at the values as mentioned in i), ii), iii) above respectively. A dump pulse of -70V, which was gated on by the falling of the detection pulse, has been applied to one of the trap's end caps. This dump pulse will drive ions out the trap to be detected by a channel electron multiplier (CEM), which is gated on by the detection pulse (leading edge). Ion signals are shown in Fig.(13 b), Fig.(14 b), and Fig.(15 b) respectively. Using the time-of-flight technique we can easily identify N⁺ and N₂⁺. The time of flight t is given by the





(b) N⁺ signal detected by channel electron multiplier at $V_o = 400V$, $f_o = 1.2MHz$, $U_o = 25V$





(b) N_2^+ signal detected by channel electron multiplier at $V_o = 255V$, $f_o = 0.65MHz$, $U_o = 20V$.





(b) N⁺ and N₂⁺ signal detected by channel electron multiplier at $V_0 = 220V$, $f_0 = 0.65MHz$, $U_0 = 20V$. equation:

$$t = s \sqrt{\frac{m}{2Vq}}$$
(4.6)

where s is the distance between center of the trap to the CEM, V is dump pulse voltage, and $\frac{m}{q}$ is mass to charge ratio of the ions. The experimental results are consistent with the calculations, i.e. when the trap parameters are set to confine both ions, according to the calculated results, we see two ion signals, but when we detune the parameters to store one kind of ion, only one ion signal appeared, and the flight time of those are consistent with Eq.(4.6). When we turn off the N₂ gas supply, no signal was observed.

In this experiment the rf frequency for the ion trap has been set at $f_o = 1.2$ MHz. The rf peak potential V_o and DC bias U_o have been scanned to find a optimum storage condition for N⁺, that is, deep spherical well, which will confine N⁺ in the relatively larger density with the maximum stability. The largest N⁺ signal was observed at $f_o = 1.2$ MHz, $V_o = 400$ V, and $U_o = 25$ V. These setting were used throughout this experiment.

The storage time of N⁺ in the trap with various N₂ pressures was determined by monitoring the decrease of the ion signal magnitude when the stored ions were driven out of the trap at progressively longer delay times. A plot of the decay curves is shown in Fig.(16). The storage time at 1×10^{-6} Torr was about 1 sec. At higher pressure, storage time decreased. Limitation of the storage time is mainly due to N⁺ loss through elastic collisions, the charge transfer process N⁺(⁵S₂) + N₂ \rightarrow N⁺₂ + N, and various quenching processes, but in all cases the storage time exceeded the metastable lifetime by at least a factor of 10.

4) Photon Detection

Photons spontaneously emitted from $N^+({}^{5}S_2)$ in the ion trap were detected by



Fig.(16) Normalized decay curves of N⁺ signal versus storage time under various N₂ pressures.

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a photomultiplier tube (PMT) and the corresponding signal was sent to a gated multichannel analyzer (MCA) where the signal was stored and displayed. The main source of noise in our experiment is the blackbody radiation from the electron gun cathode. An interference filter with 230Å bandpass centered at 2150Å was used to reduce this noise. An average dark count rate of 0.8 counts per second was typical with the electron gun heater on. Another main noise source is a fast decay ($\tau < 0.3 \text{ms}$),² which was seen and could attributed to emission in the N₂ Lyman–Birdge–Hopfield bands. This fast decay was discriminated against by delaying the detection gate with respect to the electron gun pulse. The width of the relaxation pulse was chosen to 1.6ms, which is long enough to avoid the fast decay and compensate e gun bias switching off time. The MCA, which was set at multichannel scaling mode (MCS), was then gated on by the falling edge of the relaxation pulse and started to count the photons seen by the PMT.

In this experiment, 256 channels were selected for the MCS sweep with 200 μ s per channel. So the total time for one sweep was 256 × 200 μ s = 51.2ms, which is one order of magnitude greater than the estimated N⁺(⁵S₂) lifetime. The PMT high voltage was set at 2400V for a 10⁶ ~ 10⁷ amplification factor.

5) Data Analysis

Data accumulated by the MCA is dumped to the computer and analyzed by the "DISCRETE" program. Physically, only one exponential decay is allowed for the following reasons:

- a) The trap parameters have been set such that it can only confine N^+ .
- b) The narrow band filter will only allow photons of about λ 2140Å to pass through.
- c) Non-intersystem fast decays with photons in the range of 2000Å to 2300Å are discriminated against by the 1.6ms delay. In other words, the

PMT can only see photons due to $N^{+}({}^{5}S_{2})$ to $N^{+}({}^{3}P_{1,2})$ transition.

The automatic least squares analysis of data can be represented by

$$N_{k} = \alpha_{0} + \alpha_{1} exp(-\lambda_{1} t_{k})$$
(4.7)

where

 $k = 1, 2, \dots, 256$ $t_1 = 0.2ms, \quad \delta t = 0.2ms$

and N_k is the number of photons at time t_k , whereas λ_1 is the 5S_2 decay rate.

V EXPERIMENTAL RESULTS

Six sets of measurements, taken at different N₂ pressures ranging from 1×10^{-6} Torr to 5×10^{-6} Torr, were used to obtain the final results. All other parameters were fixed. Every measurement of the radiative transition rate under certain N₂ pressure was the mean value of several runs. The standard deviation was calculated afterwards. The count rate as a function of time varies with the N₂ pressure. At 5 x 10⁻⁶Torr the count rate was 29.6 counts per second at channel 15 (recall the sweep time was 200 µs per channel) which is about one lifetime. A decay curve with good signal to noise ratio could be obtained in 6 to 14 runs with 9999 sweep cycles per run. Here one sweep cycle represents one measurement, so the data were accumulated up to 140000 measurements. A typical N⁺(⁵S₂) decay curve is shown in Fig.(17), where a nonlinear, least–squares routine with a function of the form $\alpha_0 + \alpha_1 \exp(-\lambda_1 t_k)$ was used to fit the data. The results of the fit are shown in the following table

N ₂ Pressure (torr)	Decay Rate (sec ⁻¹)
1.0 × 10 ⁻⁶	190 ± 10
1.5×10^{-6}	214 ± 11
2.0×10^{-6}	230 ± 11
3.0×10^{-6}	270 ± 13
4.0×10^{-6}	297 ± 13
5.0 × 10 ⁻⁶	342 ± 19

where three standard deviation has been chosen to ensure the results with confidence of 99%.



Fig.(17) A typical $N^{+}({}^{5}S_{2})$ decay curve. The straight line is the least-squares fit to this decay.

Different radiative transition rates with respect to the different N₂ pressures implies loss of metastable ions due to various channels of quenching, elastic collision and charge transfer. Fig.(18) shows the relation between decay rate and N₂ pressure. The slope of the decay rate versus the N₂ pressure curve gives a total ion loss rate coefficient of 1.14×10^{-9} cm³sec⁻¹ for N⁺(⁵S₂) in N₂. This curve was found to be linear in our N₂ pressure range, as expected. By extrapolating the curve to zero N₂ pressure, we get the mean N⁺(⁵S₂) decay rate of 156 ± 17 sec⁻¹.

The whole system for this experiment was carefully tested and calibrated. The MCA timing in multichannel scale mode was calibrated by 7104 oscilloscope calibrator for which the deviation at repetition rate of 1kHz was within 0.25%. The ion trap rf frequency was monitored by Ortec 878 timer/counter during the experiment. It was quite stable and the deviation was less than 0.01%. The ion trap power supplies both AC and DC were also monitored by 2235 oscilloscope to insure the ion trap operating point was right. The working condition of the e gun is also very critical to this experiment. That includes: electron gun filament heater current, electron gun positive and negative bias voltage, and electron gun electron firing stability. If the electron gun filament potential is too high, excessive electrons could go into the ion trap during electron gun "on" time. These excess electrons could partly short the end cap to the ring electrode of the ion trap thus changing the RF amplitude V_o and disturbing N^+ ion storage. On the other hand, if the filament heater current is too low, the number of electrons created by electron gun will not be enough to generate sufficient N⁺ so that the signal to noise ratio will be poor. To avoid that, we monitored the trap RF amplitude V_o as we increased the electron gun heater current so that the current could reach a threshold without changing the V_0 . The electron gun negative bias was set in between the electron gun and the end cap of ion trap to drive the electrons created by electron gun into the ion trap during the electron gun "on" time, as mentioned in III. The magnitude of the negative bias





voltage was chosen to get the maximum $N^{+}({}^{5}S_{2})$ radiation. The electron gun positive bias was set to hold the electrons during the electron gun "off" time. This voltage should be high enough so that no electrons can escape from the electron gun and get into the ion trap to produce more ions during electron gun "off" time, otherwise the new born $N^{+}({}^{5}S_{2})$ may compensate the original $N^{+}({}^{5}S_{2})$ decay, that could make the "lifetime" looks longer. To optimize this, several different positive bias voltages were set and the decay rates were measured. No difference was found in the decay rate whether +50v or +100v bias was used. To make the electron gun firing stable, we usually activate the electron gun a few hours before we do the Since we got the lifetime by extrapolating the decay rate to measurement. zero-pressure, the linearity of the pressure measurement is also very important. This has been done by comparing the Masstorr DX100 to an ion gauge. Though the ion gauge was not carefully calibrated this did not affect the decay rate, which depends only upon the linearity of the ion gauge. Compared with the systematic error, the statistical error was dominant in this work. So consequently, the measured value for the radiative decay rate of $N^{+}({}^{5}S_{2})$ is $A_{rad} = 156 \pm 17$ sec⁻¹, or, equivalently, the radiative lifetime is $\tau_{rad} = 6.4 \pm 0.7$ msec.

This result is in very good agreement with the theoretical values, but about 34% longer than the first measurement made by Knight. A reasonable explanation for this difference is that the ion trap we used is slightly different than that of Knights'. In our experiment, a meshed ion trap was used. Unwanted ions created in the ion trap during e gun "on" time, will more easily to get out of this ion trap than the ion trap with a solid wall during the "relax" time so that there would be less disturbance to the lifetime measurement. The pressure in the ion trap, after the creation of ions, tended to quickly equalize with the surroundings so that the pressure we read from the Masstorr DX100 was the same as the pressure in the ion trap during the lifetime measurement.

 $(5 \times 10^{-9}$ Torr) so the quenching from the residual gas can be neglected. We noticed that Knight used a solid wall ion trap in his measurement. This kind of trap does not have the advantages of the meshed ion trap. The residual gas pressure was 1.8×10^{-7} Torr in Knight's system so there were possibilities that some unknown quenching may happen during the lifetime measurement. Especially, if there is any contamination on the ion trap wall and they keep to vaporize in a certain rate during the lifetime measurement, these unwanted particles could then interact with the desired ions and cause additional ion loss. Such unpredictable ion loss may make the "lifetime" shorter. The pressure in the ion trap during the measurement could also be higher than the surroundings and the pressure read from the ion gauge could be different from that in the ion trap. The shift of the pressure could also affect the decay rate measurement.

VI. CONCLUSION

- 1) We created $N^+({}^5S_2)$ via electron impact with N_2 gas in a cylindrical rf-quadrepole ion trap.
- 2) N^+ has been confined in the ion trap for at least ten times longer than the $N^+({}^5S_2)$ radiative lifetime under 1 x 10-6Torr to 5 x 10-6Torr N₂ pressure.
- 3) $N^{+}({}^{5}S_{2})$ to $N^{+}({}^{3}P_{1,2})$ radiative decay has been detected. The lifetime of the $N^{+}({}^{5}S_{2})$ metastable state is 6.4 ± 0.7 msec.

Our experimental results confirmed the theory of the $N^+({}^5S_2)$ decay. It also supported the interpretation of the λ 2145Å auroral feature as being due to $N^+({}^5S_2)$ emission.

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