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NUMERICAL ANALYSIS OF THE SOLID PARTICLE SOLAR RECEIVER WITH

THE INFLUENCE OF AN AIR-JET

By

Zhuoqi Chen

Bachelor of Technology in Electrical Engineering and Automation Hangzhou Dianzi University, People's Republic of China **2006**

A thesis submitted in partial fulfillment of the requirements for the

Master of Science Degree in Mechanical Engineering Department of Mechanical Engineering Howard R. Hughes College of Engineering

> **Graduate College University of Nevada, Las Vegas December 2008**

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November 21, 2008

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Numerical Analysis of the Solid Particle Solar Receiver with the

Influence of an Air-jet

is approved in partial fulfillment of the requirements for the degree of

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ABSTRACT

Numerical Analysis of the Solid Particle Solar Receiver with the Influence of an Air-jet

by

Zhuoqi Chen

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This thesis develops and analyzes a three-dimensional computational model of a solid particle solar receiver (SPSR) for providing the heat source in a hydrogen production process using the sulfur iodine thermo chemical water splitting reaction. In this reaction, a heat input of at least $850 \, \text{°C}$ is necessary to keep high hydrogen production efficiency. Previous studies to achieve higher efficiency on a SPSR include changing particle materials, sizes, flow rates, and the geometry designs. The present study is concerned with the use of an air-jet in front of the open aperture and different operation conditions for the SPSR design optimization.

The conceptual design of the SPSR is provided by Sandia National Laboratories (SNL). There is an open aperture in front of the receiver cavity, and heat will leak to outside without any protection. Different research topics have suggested that an air-jet consisting of a transparent gas stream injected across the receiver aperture is a good method for isolating the interior from the surroundings. The main purpose of this research is to use numerical analysis to study the SPSR with the influence of an air-jet. A two-way coupled Euler-Lagrange method is applied which includes the continuity, heat, momentum exchanges between the solid and gas two-phase flows. A two-band discrete ordinate solar ray tracing model is used for the radiation interactions and heat transfer within the particle clouds, and between the cloud and the internal surface of the receiver. Different air-jet velocities are compared to evaluate the thermal performance of the receiver. Parametric studies also include varying particle size, mass flow rate, solar flux, and air-jet temperature to determine the optimal operating conditions. The temperature and velocity profiles inside the cavity are also analyzed. In all the parametric studies and thermal analysis, the SPSR with a downward air-jet velocity of 8 m/s, air-jet temperature of 300 K, and particle diameter in the range of 70-80 micron provides the best performance in the presence of a radiant flux of 920 suns. This cavity efficiency is 85%, and average exit particle temperature is 1199 K.

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NOMENCLATURE

 a_{λ} = spectral absorption coefficient *C,)=* drag coefficient C_p = specific heat capacity (J/kg-K) c_μ = constant, c_μ = 0.09 c_1 = constant, c_2 = constant, c_2 = 1.9 d_p ⁼ the particle diameter (m) *n =* the refractive index $G =$ radiation incident (W/m²) G_r = Grashof number g_i = gravity force (m/s²) $I(\vec{r},\vec{s})$ = total intensity $I_{b\lambda}$ = black body intensity $k =$ turbulent kinetic energy (m²/s²) *p =* **pressure (Pa) Pr** = Prandtl number of gas flow **Pr,** = turbulent Prandtl number $Q_{radloss}$ = radiation heat loss (W)

 $\mathbf X$

 Q_{inc} = total incident solar energy (W)

 Q_{cavloss} = convection heat loss (W)

 \overline{Q}_p = integrated absorption energy gain of the particles (W)

 \overline{Q}_s = the total incident solar energy (W)

 \vec{r} = position vector

S = modulus of the mean rate of strain tensor

 S_i^M = momentum per unit volume added to the gas phase by the

presence of particles

 s_i^{τ} =energy per unit volume added to the gas phase by the

presence of particles.

 \vec{s} = direction vector

 $u_{i,j}$ = velocity of air flow (m/s)

 U_i = mean velocity of air flow (m/s)

 u'_{i} = fluctuant velocity of air flow (m/s)

 $u_{p,i}$ =particle velocity (m/s)

 $Re =$ Reynolds number

 Re_p = relative particle Reynolds number

 T = temperature (K)

 T_{P-exii} = average exit particle temperature (K)

 T_p = particle temperature (K)

 \overline{T} = mean temperature (K)

 $T =$ fluctuant temperature (K)

 T_R = the radiation temperature (K)

 $t =$ time (s)

 x, y, z = coordinates

 $x_{i,j}$ = coordinates

Greek symbols α = thermal conductivity (W/m·K) λ = wavelength (m) σ = Stefan-Boltzmann constant (W/m²K⁴) σ_k = constant, σ_k =1.0 σ_{ε} = constant, σ_{ε} =1.2 σ_{θ} = constant, σ_{θ} =0.9 σ_s = scattering coefficient (1/m) ε = turbulence dissipation (m²/s³) ε_p = particle emissivity (1/m) η = cavity efficiency ρ = air density (kg/m³) ρ_p = particle density (kg/m³) μ = viscosity (kg/(m·s)) μ_{τ} = turbulent viscosity (kg/(m·s)) $v =$ kinetic viscosity (m²/s)

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CHAPTER 1

INTRODUCTION

1.1 Introduction

Renewable energy like solar, wind, and geothermal resources will become the most important energy supplies for both domestic and world wide consumption in the near future, to eliminate rolling black outs and the increasing cost of energy. People will focus more on the utility part of transforming renewable energy to the form of energies that can be applied to the daily human being activity. The drawbacks to fossil fuel use include limited reserves, and carbon dioxide emissions, which is a greenhouse gas responsible for global warming. Fossil fuel combustion is also responsible for environmental pollution. Hence, the critical part of transforming the energy from a renewable source is to ensure there is no environmental waste or cost for future generations.

Hydrogen, a promising and clean energy carrier, could potentially replace the use of fossil fuels in the transportation sector. The demand for hydrogen is ever growing, with major applications in refinery use and as a component in synthesis gas for manufacture of methanol, ammonia and gas to liquids (GTE) products. Another application for hydrogen in the future will be the fuel cell and internal hydrogen combustion engine for the automotive sector, which could result in a large demand of hydrogen. Currently, no environmentally attractive, large-scale, low-cost and high-

efficiency hydrogen production process is available for commercialization [1]. Hydrogen is an important reactant and an energy carrier that can be manufactured from a range of energy sources such as fossil fuels, biofuels, renewable energy, and nuclear energy via electricity [2], Today's hydrogen production is mainly based on fossil fuels and most specifically on natural gas [3]. There are no environmental benefits if the hydrogen is derived from natural gas or fossil fuel reforming, because of the emissions of $CO₂$ from these production routes.

The use of wind, hydropower and solar thermal energy for the production of hydrogen are the most environmental friendly methods. The solar-driven water-splitting thermochemical cycles may constitute the optimized choice for hydrogen production. Many cycles can be selected for the production of hydrogen from splitting water, and most of them need a reaction temperature in the range of 850-2000 $^{\circ}$ C. The sulfur-iodine and hybrid sulfur processes are two examples of the hydrogen production process having a maximum temperature of around 850 °C [4], [5] which can be supplied by nuclear or solar energy. Reactions that need more energy input, like the zinc-oxide [6] process, require heat input at temperature of 2000 °C. The only energy supply which can reach this high temperature is from solar energy.

As energy prices have increased, engineers have focused more on the overall cost of fuel. One benefit for using renewable energies is their unlimited resouree, so operating cost becomes the most important part of the total energy production process. There are several designs for producing hydrogen by using renewable energy as the heat source, and one of them is through using solar energy. Hydrogen eould be produeed by thermoehemieal water-splitting, a ehemical proeess that eonsists of the multistep

 $\overline{2}$

decomposition of water. Water and heat are the inputs, and hydrogen and oxygen are the only outputs. The other chemieals and reagents are reeyeled in a elosed eyele. The heat ean be applied using nuelear energy, but it is much more enviromnentally aceeptable if the heat souree is solar energy. One problem with using solar energy is although the solar radiation is a high quality energy souree, its power density at the earth's surfaee makes it diffieult to extract work and achieve reasonable temperatures in a working medium. Therefore, solar thermal energy must be collected by concentrated methods.

Coneentrated solar energy ean be eolleeted by a solar eentral reeeiver system whieh uses mirrors (large seale of heliostat field) to refleet and foeus sunlight onto a receiver located on top of a solar tower. In this manner the collected sunlight can reach hundreds of suns at the surface of the solar receiver. The concentrated solar energy is eapable of attaining the temperature requirement for the thermoehemieal reaetion eyele at 2000 °C.

A working medium is needed to transport the energy for the reactions that need a heat source. Previous studies on the working medium in a solar central receiver have been on gases or liquids whieh need to be transported by tubes or pipes. There are working demonstrations of systems using liquid as the heat medium like the water/steam eentral reeeiver system, loeated near Barstow, California, which produces eleetrieal energy [7]. The other common working fluids are molten salts, molten metals, and air.

The term solid particle solar receiver is another concept of a heat energy transfer medium with great interest. Sand-size refraetory particles fall down freely inside a solar receiver to form a curtain that directly absorbs the solar insolation. The advantages of using a solid particle reeeiver over the systems of gas/fluid mediums are; (1) the solid

particles ean absorb the reflected and eoneentrated solar energy direetly, eliminating the pipes or tubes used to transport the fluids, (2) solid material might reach higher temperature, (3) solid partieles can be a heat storage medium, eliminating the need for extra elements in the storage system [8],

The solid particle solar receiver turns out to be an attraetive ehoice as the heat souree for hydrogen production. Most solid particle reeeivers have an open aperture in front of the cavity where the concentrated solar light can pass through, but a large amount of heat might be lost from there by conveetion and radiation. The reasons for abandoning the transparent window to eover the open aperture are that materials that endure high temperature are not readily available, and the solid partieles are very light and small and will stick on the window when spreading inside the cavity, reducing the window transpareney. In this thesis, an air-jet eonsisting of a transparent gas stream injeeted across the aperture to isolate the interior from the outside was used. The benefit is the elimination of eonveetion heat loss [9].

An air-jet is used routinely as a thermal barrier for warehouse doorways and department store entranees. For the aperture of the SPSR, the eommereial air-jet is applicable for the whole system in the applieation methods, and the cavity efficiency will increase by the insulation of the interior from the ambient environment. On the other hand, the fluid dynamie behavior is very eomplieated inside the whole system, especially when coupled with the air-jet and characteristics of the particles.

In the engineering world, eomputational fluid dynamies (CFD) is a useful numerieal tool for the analysis of problems whieh involve fluid flow and heat transfer. High requirements for computer performance are needed during the calculation process.

The most fundamental consideration in CFD is treating the continuous fluid in a discretized way, by dividing the domain into small meshes or grids. Then, it is neeessary to apply meshes into a suitable algorithm to solve the equations of continuity, motion, energy, and species. Compared to experimental research, CFD analysis provides a simulation environment for each case without wasting material and time, which is expensive and unpredictable. Continuous flow is complicated by unstable flow patterns, so different kinds of conditions and models ean be applied to solve these problems using CFD analysis. The aerodynamic behavior inside the solid particle solar receiver can be evaluated by this computational tool.

1.2 Literature Review

The solid partiele solar receiver studies started about 20 years ago. There are plenty of researchers considering methods of hydrogen production by splitting water or other chemical reaction routines. These kinds of processes require very high temperature input as the heat source to realize the optimized design for the whole hydrogen production system. Solar energy is an unpolluted and high temperature heat supply. For engineers, the design of colleeting the energy is very important. In some designs, the heat absorbing and reaction progress occur in the same place [3-6]. In this present design, the SPSR is only used as a heat absorbing deviee. Solid gas two-phase flows can happen in several design domains, and will be discussed in this review. Also, the properties of materials for solar absorptions were experimentally studied by SNL. In the simulation section, different modeling results for the radiation, heat convection and turbulence flow from other researchers are reviewed.

1.2.1 Experimental Investigations of the Solar Thermal Hydrogen Produetion and Solar Reeeiver Design

Different designs of eoneentrated solid particle receivers were used as solar energy colleetors or both solar energy colleetors and thermoehemieal reactors. Inside this receiver cavity, solid gas two-phase flows are analyzed. The design of a concentrated solid particle receiver depends on the temperature requirement of the reaction cycle and the desired system working efficiency.

Life cycle assessment (LCA) of hydrogen production was first considered by Koroneos et al. in [3]. In his research, LCA is a powerful tool to help evaluate the impact from a process or from production and use of a product. It consists of goal definition and scoping which defines the product, process or activity; inventory analysis which identifies material usage and environmental releases; impact analysis which assesses the human and ecological effects of energy, water and material usage; and last interpretation, which evaluates the results of each analysis. The research compared hydrogen produetion by natural gas, renewable energy, electrolysis, and fuel. The use of wind, hydropower, and solar thermal energy were proven to be the most environmentally friendly and efficient methods. And the future of renewable hydrogen energy depends strongly on reduced costs for renewable energy production.

The methods of collecting renewable energy such as solar thermal energy for hydrogen production are concluded by some scientists. Stephane summarized a database of water-splitting thermoehemieal cycles [1]. All seleetions and evaluations of promising thermochemical reaction cycles were performed in the temperature range of 850-2000 $^{\circ}$ C. Process temperature, process eomplexity, and energy analysis of thermoehemieal cycles

were used as the standards for efficiency, and the sulfur-iodine (S-I) eyele was the eyele with the highest reported efficiency (52%). It was suggested, that with process improvements, it was possible to inerease the effieieney and lower the eapital eost [10] [11]. The S-1 eyele consists of three reaetion steps [12]:

$$
H_2SO_4(g) = SO_2(g) + H_2O(g) + 0.5O_2(g) \qquad 850 °C \qquad (1.1)
$$

$$
I_2(l) + SO_2(aq) + 2H_2O(l) = 2HI(l) + H_2SO_4(aq) \qquad 120 \,^{\circ}\text{C}
$$
 (1.2)

$$
2\text{HI (l)} = I_2 \text{ (l)} + H_2 \text{ (g)} \qquad \qquad 450 \text{ °C} \tag{1.3}
$$

The cycle can be separated into four sections:

I. Acid production and separation with oxygen purification,

II. Sulfuric acid concentration and decomposition,

III. Hydroiodic acid (HI) concentration,

IV. HI decomposition and H_2 purification

Section II of the cycle is an endothermic reaction driven by the heat input from a high temperature souree. The utilization of nuelear power and solar energy ean be considered as the temperature heat source in this step [13]. William A. Summers has researched nuclear hydrogen production based on the hybrid sulfur thermochemical process [5]. This requires two sulfur eyeles - the sulfur-iodine and the hybrid sulfur. The second thermal decomposition requires a heat input of 800-900 °C, which is provided by nuclear reactors. Process design studies suggest that a net thermal efficiency of over 50% is possible with this hydrogen produetion eyele. The heat input required to drive the solar thermal reaction was provided by the solar reaetor as a eavity reeeiver. Several designs for these kinds of thermal reaetions involving gas and solids are introdueed in the following paragraphs.

The solid particle solar receiver has been designed as the main component for the gas-solid thermoehemieal reaetions in Steinfeld's research [14]. In his work, the reaction of interest is characterized by the equation;

$$
CaCO3 = CaO + CO2
$$
 1170 K (1.4)

This reaction will go spontaneously to the right at temperature above 1170 K and 1 atm. As the reaction system is complicated by the heterogeneous reactants and products which need to be removed, separated, and fed, the feature of both the solar receiver and reactor need to be combined. Incoming concentrated solar radiation enters the cavity without any window protection, because windows for solar receivers are usually expensive, brittle, require careful mounting, but often fail to withstand the high solar fluxes and high temperature, redueing the solar energy absorption efficiency. The experimental set-up is done, and the peak flux eoneentration ratio observed was 1410 suns (1 sun =1 kW/m²). The reactor in the solar furnace with the $CaCO₃$ decomposition reaction reached 1300 K and obtained high degree of calcination. The energy absorption efficiency, based on the energy incident on the receiver aperture, was 43 % [14].

The production of zinc by reduction of zinc oxide and refining of $CH₄$ is characterized by the high energy consumption, using solar thermal energy as energy source to eause the following reaction:

$$
ZnO(s) + CH_4(g) = Zn(g) + 2H_2(g) + CO(g) \qquad 1200 K \qquad (1.5)
$$

$$
Zn(g) + H_2O = ZnO(s) + H_2(g)
$$
 (1.6)

This reaction is studied in a solar receiver consisting of a fluid-bed reactor and a compound of parabolic concentrators. The ZnO partieles are directly irradiated, fluidized in CH4, and aeted as heat absorbers and reactants. The produced Zn vapor was trapped in a condenser, and synthesis gas containing a 2:1 mixture of H_2 and CO is collected, without diseharging greenhouse gases and other pollutants. This is a relatively environmentally clean path for either recycling Zn -air or producing H_2 in a water-splitting seheme [15]. The production of Zn from ZnO in high temperature ean be used to split water to hydrogen in an exothermic reaction [16], and the by-product of ZnO can be recycled to the solar process.

Another study analyzed the radiation heat transfer in a solar chemical reactor for the eo-production of hydrogen and earbon by thermal decomposition of CH4.

$$
CH_4(g) = C(s) + 2H_2
$$
 800-1500 K (1.7)

The solar chemical reactor features a vortex flow of CH₄ with the participating of solid carbon particles to serve both as the heat absorber and the reaetion eatalyst. Hirsch [17] developed a three-dimensional reaetor model to prediet the energy, temperature and ehemieal reaction inside the solid particle cavity. Different parameter inputs like particle size and solar loading were validated to improve the overall reaction efficiency.

1.2.2 Experimental and Numerical Studies of Gas Flow with Solid Interaetion

Computational fluid dynamies (CFD) was employed in the design and optimization of a high temperature central receiver or reactor, which combined the gassolid flows inside. Experimental works aeeompany the simulation results to identify the eharacteristies of the two-phase flows.

The fluid flow, partiele transport, and heat transfer of a high-temperature solar ehemieal reaetor was studied by Meier et al. [18]. The reaetor eonsists of a eylindrieal eavity and an open aperture through whieh eoneentrated solar radiation enters. The partieles (reaetants) impinge on a cone and are conveyed in a swirling air stream. A general purpose CFD eode was used to simulate the fluid flow and heat transfer in the partiele-eloud reactor. The governing Navier-Stokes equations were solved by a 'hybrid' differeneing seheme. The semi-implieit method for pressure-linked equations eonsistent (SIMPLEC) velocity-pressure eoupling algorithm is used. Sinee the eonventional $K - \varepsilon$ turbulence model always fails to predict swirling flow correctly, a renormalization group modification of the standard high Reynolds number version of the κ - ε turbulence model was applied. The motion of partieles in the air flow was modeled using a diserete trajectory (Lagrangian) approach. An iterative proeess was needed to adjust the eoupling between the particle transport and the fluid flow. Heat transfer and particle flow pattern were presented in this researeh to evaluate the performance and validate the experimental results.

In another work, Meier [19] used the thermal heat from a falling particle receiver/reactor which exposed to concentrated sunlight to decompose limestone **(CaCOs).** In this ease, eoneentrated solar radiation is absorbed direetly by a eurtain of free-falling solid partieles that were heated to temperatures in **excess** of 1200 **K** over approximate 5 m fall in the presence of a radiant flux of less than 1 MW/m^2 . The simple schematic of the solar falling reaetor is shown in Figure 1.1. For this study, the general purpose CFD code CFX-4 was used, the gas/particle flow and the convective heat transfer in the solar falling partiele reaetor (FPR) were modeled within the CFX-4 solver, and partiele heating by radiation was calculated in the CFX-RADIATION model using a Monte Carlo method. Input parameters and results from CFD simulations of the FPR are compared. The model was verified with experimental data and an improved design of the FPR was mentioned there.

Figure 1.1 Simplified schematics of the solar falling particle reactor (FRP) [19]

A detailed recent study about the two-phase flow gave insight about the radiation characteristics of particles, which are approximated using seeded particles of variable size and shape [20]. It proved that radius of limestone less than 100 micron is inefficient in absorbing solar energy. In present study of SPSR, the solid particle used has a higher heat capacity (Cp= 1085 J/kg-K) than limestone (Cp= 840 J/kg-K). So particles with radius less than 100 micron still can perform with high efficiency in this study of SPSR. Different from previous studies, molecular particle collisions were considered in the energy transfer. The receiver simulations were performed by the commercial CFD code PHOENICS, coupled with two radiation transfer solvers. A swirl model was added to the standard PHOENICS code to simulate the tangential flow in the receiver. And a Monte Carlo ray-tracing method was used to introduce the irradiation incident on the aperture plane. The receiver simulations showed good agreement with the wall temperature distribution that measured in experiment. Due to the limitation of the simulation code and model, gas exit temperature was significantly lower than the measured value.

There are other models for the simulation of two-phase flow ereated in the 1970- 80's, which were based on one-dimensional or two-dimensional geometries and were relatively simple. In Crowe et al.'s [21] research of gas-droplet flows, a particle-source-in cell (PSI-CELL) model was used to illustrate the capability of the model to treat the complex phenomena associated with multiphase flows. Raithby introduced a "finitevolume" method to predict radiant heat transfer in enclosure with participating media [22]. This method conceptually used the same nonorthogonal girds to compute fluid flow and heat transfer. The results have been benchmarked and have shown good accuracy.

1.2.3 Numerical Method for Turbulence Flow

In engineering practice, flow in three dimensions liking pipes, cavities and wakes, will become complicated when the Reynolds number is higher than a certain value. This is observed as turbulent flow, a chaotic and random state of motion that develops in which the velocity and pressure change continuously with time and within substantial regions of flow [23].

A turbulence model can be used by calculating the mean flow, so that less variety of flow problems can be calculated. A turbulence model must be applicable to the general CFD code and be accurate, simple, and economical to run. The most common turbulence models are classified in Table l.I. This table is divided into classical models and large eddy models. In this review, much more attention is paid to the numerical and experimental variability of the two-equation $\kappa - \varepsilon$ model.

Table 1.1 Turbulence models [23]

Classical Models	Based on Time-Averaged Reynolds Equations			
	zero equation model - mixing length model two-equation model - $\kappa - \varepsilon$ model 2. 3. Reynolds stress equation model 4. algebraic stress model			
Large Eddy Simulation	Based on Space – filtered Equations			

Jones used a two-equation model for the prediction of turbulence flow early in the 1970's [24]. In the two-equation turbulence model, the turbulent viscosity is determined from the solution of transport equations for the turbulence kinetic energy and the energy dissipation rate. In high Reynolds number flow, the turbulence energy and energy dissipation can be written as:

Turbulence energy:

$$
\rho \frac{D\kappa}{Dt} = \frac{\partial}{\partial y} \left(\frac{\mu_r \partial \kappa}{\sigma_k \partial y} \right) + \mu_r \left(\frac{\partial u}{\partial y} \right)^2 - \rho \varepsilon
$$
\n(1.8)

Energy dissipation:

$$
\rho \frac{D\varepsilon}{Dt} = \frac{\partial}{\partial y} \left(\frac{\mu_r \partial \varepsilon}{\sigma_\varepsilon \partial y} \right) + c_1 \frac{\varepsilon}{\kappa} \mu_r \left(\frac{\partial \mu}{\partial y} \right)^2 - c_2 \frac{\rho \varepsilon^2}{\kappa}
$$
(1.9)

The turbulence viscosity is defined by $\mu_r = c_\mu \rho \kappa^2 / \varepsilon$. The terms c_μ , c_1 , c_2 , σ_κ , and σ_{ε} are empirical constants, the values of which are defined in the high Reynolds number form of the $\kappa - \varepsilon$ model of turbulence.

Later on, a model for numerical simulation of buoyant, turbulent flow was developed by Humphrey et al. [25] [26]. The first part corresponds to free convection along a heated vertical flat plate, and the second part extends to predict steady free and mixed convection flows of air in a strongly heated cavity of arbitrary rectangular crosssection and orientation. Both cases consider the numerical simulation of more complex, buoyancy-affected, turbulent flows. Two low Reynolds number turbulence model formulations had been developed for predicting wall-bounded, variable property, free convection flows. The KEM model (commonly referred to as the $\kappa - \varepsilon$ model) relates turbulent fluxes to eddy viscosities via a generalized Boussinesq hypothesis. The latter approach involves partial differential equations for the turbulence corrections. Truncation of these transport equations, obtained by neglecting convection and diffusion terms, yields a system of algebraic equations relating the turbulent fluxes to known or calculable flow quantities; the terminology algebraic stress model (ASM) is applied to describe the method. In the second part of these studies, the characteristics of the flow depend on the cavity aspect ratio, a/b , the inclination angle, α , and the Grashof number, Gr. Prediction of the Nusselt number *Nu* for the mixed convection flow is important when considering the influence of buoyant force, and it is characterized by Re^2/Gr . For both free and mixed convections, the predicted flow patterns are in good qualitative agreement with the flow profiles in experiments.

Ince **et** al. [27] **give another example** of **the calculation** of buoyancy-driven turbulent flows in rectangular enclosures under the two cases of different aspect ratios, 30:1 and 5:1. A version of the Jones-Launder low-Reynolds-number $\kappa - \varepsilon$ model with the modified coefficient was used. The turbulent stresses and heat fluxes are given by:

$$
\overline{\rho u_i u_j} = \frac{2}{3} \delta_{ij} \rho \kappa - \mu_r \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)
$$
(1.10)

$$
\rho \overline{u_j T} = -\frac{\mu_r}{\sigma_\theta} \frac{\partial T}{\partial x_j} \tag{1.11}
$$

Here $\mu_T = C_\mu \rho \kappa^2 / \varepsilon$, the turbulence energy κ and ε part of energy dissipation rate terms that are obtained from their own transport equations. The numerical results lead to a satisfactory agreement with reported experimental data.

A new $\kappa - \varepsilon$ model consists of a new dissipation rate equation and a new realizable eddy viscosity formulation is used for high Reynolds numbers turbulent flows [28]. Results of this model are compared with available experimental data and the results from a standard $\kappa - \varepsilon$ model, and it presents a significant improvement in the following types of flows:

- i. Rotating homogeneous shear flows;
- ii. Boundary-free shear flows including a mixed layer, planar and round jets;
- iii. A channel flow, and flat plate boundary layers with and without pressure gradient;
- iv. Backward facing step separated flows.

1.2.4 Experimental and Numerical Analysis of the 2D and 3D SPSR without Considering the Influence of an Air-jet

The research and development of a SPSR was conducted within the U.S Department of Energy (DOE) as a part of the Solar Hydrogen Generation Research program. Experiments on the design, and material testing, have been conducted by SNL since the 1980's. In initial research, Evans et al. [29] mentioned numerical modeling of a SPSR. They studied the flow of air and particles with combined heat transfer inside a solar heated, open eavity containing a falling particle eurtain. Two-way momentum and thermal eoupling between the partieles and the air was ineluded in the analysis, along with the effeets of radiative transport within the particles, between the heated gas and particle surface. This was a two-dimensional steady flow design, and the PSI-Cell (partiele source in cell) eomputer eode was used to deseribe the gas-particle interaction. The radiative heat transfer characteristics within the particle clouds were obtained using a diserete ordinate method. From the experimental work, particles were diluted as they were falling down, and the volume fraction of particles was small enough to ignore the particle-particle collisions. Pressure-velocity coupling was determined by using the revised procedure of semi-implicit method for pressure-linked equations (SIMPLER), and differential equations of describing turbulent energy κ and dissipation ε were solved. In the analysis of the aerodynamic behavior of the SPSR, the authors considered the interaction of particle-particle radiation, partiele-wall radiation, particle-air eonveetion and air-wall convection. The properties like scattering, thermal emission, and wavelength dependence of the particles are aecounted for in the radiation model calculation. Geometry of the two-dimensional design was a 6.7 m by 5.0 m cavity. Nominal parameters in this research were incident solar energy of 920 suns, particle diameter of 650 microns, particle mass flow rate of 5.4 kg/s, and initial particle temperature and downward velocity of 293 K and 0.3 m/s. In the parametric studies of the receiver behavior, the mass flow rate, particle diameter, particle infrared scattering albedo, and absorptivity are varied independently to determine their efficiency and particle exit temperature.

The results identified that smaller partieles can provide better performance, as shown in Figure 1.2, due to a greater optical thickness *and* longer residence time. Increasing the particle mass flow results in an increase in cavity efficiency with a decrease in partiele temperature, as shown in Figure 1.3. Solar flux intensity is a value that depends on the heliostat area and receiver aperture dimension of the eavity. In this two-dimensional design, cavity efficiency, average exit temperature of the particles, and distribution of the thermal losses under the influence of different solar flux data are compared in Table 1.2.

Figure 1.2 Variation of cavity efficiency and average exit temperature of particles as a function of particle size [29]

Figure 1.3 Variation of cavity efficiency and average exit temperature of particles as a function of mass flow rate of particles [29]

Table 1.2 Cavity efficiency, average exit temperature of particles, and distribution of heat loss as a function of heliostat field size and cavity size [29]

Field Size	Average		$T_{\text{p-exit}}$	Q_{radloss}	$Q_{\textit{cnvloss}}$
(m ²)	Flux (suns)	η	(K)	Q_{inc}	Q_{inc}
10 ⁴	790	0.59	930	0.29	0.11
10 ⁴	1280	0.70	1032	0.22	0.08
10 ⁵	650	0.59	932	0.11	0.28
10 ⁵	650	0.64	973	0.09	0.26
10 ⁵	920	0.70	1032	0.10	0.19
10 ⁵	1530	0.86	1175	0.07	0.07

Later, the aerodynamic and thermal interaction between gas flow and solid particles in a detailed three-dimensional design was investigated by Chen et al. [30]. Improved simulation tools for the modeling were applied in this research. The working media for the solid particles was bauxite and includes \sim 7% iron oxide. Realizable $K - \varepsilon$ turbulence model and solar ray tracing algorithm were used to evaluate the model. Figure 1.4 shows parametric studies on the performance of the SPSR with and without bottom opening, and less convective loss between particle and entrained air flow can be achieved without a bottom opening. Compared with larger diameter particles, particles with smaller size are easier to heat and can achieve a higher temperature.

Figure 1.4 Cavity efficiency as a function of particle size with and without bottom opening [30]

Increasing of the particle mass flow rate can also improve the cavity efficiency evidently. The performance of the SPSR can be enhanced by assuming that there is no bottom opening, and in practical applications, it can be realized by applying a hopper to collect these falling particles.

1.2.5 Experimental Research of Material Selection for Solid Particles

Particle material selection is the most important topic in the technical studies. The design of the SPSR is based on the understanding the behavior of the free falling particles, such as the agglomeration, fracture and optical properties of the particles. Several characteristics of different materials are compared by Abdelrahman et al. [31], and Hruby [32] [33]. They also mentioned the different methods that were used to analyze the behaviors of these particles.

The characteristics of the particles were first investigated in the study of solidgas-suspensions, which is used for direct absorption of concentrated solar radiation [31]. This study makes the choice of appropriate type of particles for any further studies on the solar direct absorption particles. It considers the appropriate diameter in gases, and also absorptivity and emmissivity in the IR range were investigated. Early in 1908, a classic equation for the optical behavior of spheres of any size or material that subjected to radiation was derived [34]. And this theory was used for finding out the important factors influencing the particle numerical calculation. For a given concentration and optical depth, the absorptivity depends on its complex index of refraction in a specific wavelength. This study showed that graphite is a suitable material from these optical properties and availability.

In Hruby's [32] technical feasibility studies of the solid particle behavior, the author points out that particle velocity and volume fraction are the two most important parameters for characterizing the aerodynamic behavior of a free-falling particle cloud. And in the material selection, agglomeration, fracture resistance and optical properties of alumina, silica, silicon carbide and zircon were compared. Agglomeration investigations indicated that alumina, silica, and zircon have the best performance of pure ceramic material for use between 1000 and 1200 °C. Impact-induced fracture was not a problem for the small, round particles. Bulk solar absorptivity investigation identified that the candidate material which performs best in agglomerating had low solar absorptivities. As a result, more materials needed to be tested. Two particle materials were chosen to examine under this situation. They are:

1. Master Bead tm is an alumina based particle,</sup>

2. Zirproptm is a zircon-silica (zirconia) based particle.

The results indicate that Master Beadtm is a candidate as a solid particle material at temperature up to 1100 °C. At temperatures above 1000 °C, an applied pressure will cause it to agglomerate. Zirproptm does not agglomerate at temperatures up to 1200 °C under applied pressure. However its absorptivity decreases as the temperature increases.

The working media for a SPSR is commercially available for purchase from CarboHSP according to recent research [13]. A schematic design of the drop test platform shows the material's properties with sphericity of 0.9, roundness of 0.9, and diameter of 697 micron. The solid volume fraction is measured by a circuit. Blacklit images are taken to evaluate the opacity of the curtain. From this study of particles, the particle-particle
collision can be neglected in future calculations, because the solid volume fraction is very small [35].

1.3 Research Objective

A conceptual design of a SPSR and on-sun test has been proposed by SNL. The objective of this research is to create a detailed three-dimensional model for a SPSR with the influence of an air-jet. The aerodynamic behavior of the particles and thermal interactions which include particle-particle radiation, particle-wall radiation, particle-air convection, and air-wall convection are investigated in this research. The parametric studies are aimed to evaluate the optimal performance of the SPSR. Numerical results obtained from the numerical model include the fluid flow pattern, heat transfer behavior inside the cavity, average exit particle temperature and cavity efficiency. Conclusions are drawn from these results and intended for use in further studies to improve the design of the SPSR.

1.4 Outline of Thesis

A three-dimensional conceptual design of the SPSR and details of the simulation models are discussed in this work. Chapter 2 gives a description about the problem and the numerical modeling tools which are applied in the conceptual design. The whole system is described on a conceptual basis, which is also introduced in Chapter 2. In this research, all the numerical methods and algorithms are introduced in Chapter 3. Several assumptions are made to analog the working environmental of the SPSR. In Chapter 4, the results of model of the SPSR with the influence of an air-jet are discussed, and different parametric studies are presented to find an optimized design for the model. At last, in Chapter 5, conclusions and recommendations from this research are pointed out to provide suggestions on the design work of hydrogen production from a SPSR energy source system.

CHAPTER 2

DESCRIPTION OF THE PROBLEM AND GEOMETRY

SNL provided the conceptual design of a SPSR and schematic of this energy source supply and collecting system. The experimental research of the whole system was also developed by SNL. Figure 2.1 [32] shows the solid particle receiver system design, which contains the particle lift system, hot particle storage system, direct contact heat exchanger, and lower temperature solid particle storage system. The main focus of this thesis is on the design of the solid particle receiver part as shown in Figure 2.2 [32].

Figure 2.1 Schematic of a solid particle solar receiver system [32]

Other components of the system are considered less important than the SPSR, but are with particular interests to improve the efficiency of the whole system. The particle lift system needs to move the solid particles from ground level to the top of the receiver, and promise the particle initial temperature won't below requirements before falling down from the top. Agglomeration and fracture should be prevented when the high temperature particles rest in the hot tank. These depend on both the particle material selection and design of the storage component. After particles are transported through the heat exchanger, a temperature of minimum requirement for hydrogen production reaction must be confirmed. In the hydrogen production system, a large amount of heat transfers through each component. A reasonable design can reduce the loss of heat energy and get a better performance from the SPSR.

Figure 2.2 Conceptual design of a solid particle solar receiver [32]

Aerodynamic and thermal behaviors of the free-falling particle cloud were studied, aimed at evaluating the performance of the SPSR. Experimental work of particle falling behavior concluded that particle volume fraction was small, less than 1% [29]. Under this condition, particle-particle collisions can be neglected, because particles were diluted in free fall.

2.1 Problem Description

Figure 2.3 shows three-dimensional schematic illustration of a SPSR with heated particles under the influence of an air-jet. The particles are pouring down from the top of the cavity to form a curtain which can absorb the solar irradiation directly. Air entrainment was shown to be important in experimental work of Evans et al. [29] and Koroneos et al. [3], in which the velocity of the particles was significantly higher on the terminal values, corresponding to an isolated particle falling in a quiescent environment. A buoyant force is generated because air beside the particle curtain is warmed up by the heated particles. This results in an increase in the particle residence time, which helps each particle to remain within the radiant flux field for a longer period.

In the case without an air-jet, as shown in Figure 2.4, warm air leaks from the top, and cold air sucks from downside through front aperture. If an air-jet is blowing vertically downward and cross the aperture, as shown in Figure 2.3, the inside gas-solid flow pattern and heat energy of cavity can be considered as being well protected. There exists an optimizing velocity working value of the air-jet, smaller than which heat generated by the particles can also leak from the bottom of the aperture, and larger than that value will become a waste of energy input.

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Figure 2.3 Three-dimensional schematic illustration of aerodynamic and thermal behavior in a SPSR with an air-jet

Figure 2.4 Two-dimensional schematic illustration of aerodynamic and thermal behavior in a SPSR without an air-jet [32]

2.2 Conceptual Design of Solid Partiele Solar Receiver

The SPSR geometry and dimensions are shown in Figure 2.5. This is the conceptual design for the SPSR that may be constructed by SNL for experimental verification. Our research objective is to prove numerically that the design can achieve a high exit particle temperature (>850 °C), which is needed for the heat supply of the sulfur-iodine thermoehemieal process. The geometry of the SPSR consists of a rectangular cavity that is 2 m in width, 1.58 m in depth and 3 m in height. On the top of the eavity, an opening of 0.25 m by 1.5 m is used for introduction of the particles. An area of 1.5 m by 1.5 m square is the open aperture in front of cavity, from the top of which an air-jet will blow downwards.

Figure 2.5 Conceptual design of the SPSR for the computational fluid dynamics (CFD) study (unit: m)

CHAPTER 3

NUMERICAL METHOD AND ALGORITHM

This chapter introduces the numerical methods and algorithm that are used in evaluating the performance of the solid particle solar receiver with the influence of an airjet. Assumptions and system working boundary conditions are described, and the governing equations and simulation models are introduced in this chapter.

In this research, a two-way coupled Euler-Lagrange method is implemented. The momentum and heat transfer between the particles and gas are considered. The governing equations used to express gas-solid thermal aerodynamics are solved in the Cartesian coordinate system with a control-volume finite difference method which is introduced by Patankar [36]. The commercial software Fluent is used to simulate the performance of solid particle solar receiver.

3.1 Numerical Model Assumptions

Several assumptions in the numerical model are made to investigate the performance of SPSR. For the simplicity of describing SPSR, a few assumptions may be too ideal, but they provide a good standard to analyze the important factors which influence performance. Numerical model assumptions are listed below.

1. Solar flux is in the direction normal to the front open aperture with a fixed value.

- 2. The size of the particles is assumed to be the same, and with the same spherical shape.
- 3. Air is assumed to be incompressible and follow the ideal gas law.
- 4. The particle properties are constant, no agglomeration occurs and no fracture in high temperature occurs.
- 5. The environmental influence on the SPSR like wind, temperature and humidity are ignored in this study.
- 6. The wall of SPSR is assumed to be adiabatic.

3.2 Governing Equations and Numerical Model

3.2.1 Air Flow Equation

Inside the SPSR, falling particles can achieve a very high temperature by absorbing radiation heat energy. At the same time, the air surrounding the falling particles is also heated up to form a complicated gas-solid flow pattern, which contains rotation and recirculation. The Reynolds number can reach 1.6×10^5 in this problem. As this is a typical turbulent flow, the velocities and temperatures items in the calculation equations contain a mean and a fluctuating part, $u_j = U_j + u_j$ for velocity and $T = \overline{T} + T$ for temperature. The time-averaged partial differential governing equations for describing the gas phase mass, momentum, and energy can be expressed as:

Continuity:

$$
\frac{\partial \rho U_j}{\partial x_j} = 0 \tag{3.1}
$$

Mómentum equation:

$$
\frac{\partial}{\partial x_j} \left(\rho U_i U_j \right) = -\frac{\partial P}{\partial x_i} + \frac{\partial}{\partial x_j} \left[\mu \left(\frac{\partial U_i}{\partial x_j} + \frac{\partial U_j}{\partial x_i} \right) - \rho \overline{u_i u_j} \right] + S_i^M - \rho g_i \tag{3.2}
$$

Here ρ is the density, μ is the viscosity, P_1 is the pressure, and S_i^M is the momentum per unit volume added to the gas flow by the influence of particles. The turbulence stress is described as:

$$
\overline{\rho u_i u_j} = \frac{2}{3} \rho \delta_{ij} k - \mu_r \left(\frac{\partial U_i}{\partial x_j} + \frac{\partial U_j}{\partial x_i} \right)
$$
(3.3)

 $\mu_{\tau} = \rho C_{\mu} k^2 / \varepsilon$ is the turbulence viscosity, where $C_{\mu} = 0.09$ and ε is the turbulence dissipation. And *k* is the kinetic energy, δ_{ij} is the Kronecker delta function, $\delta_{ij} = 1$ when $i = j$ and zero when $i \neq j$.

Energy equation:

$$
\frac{\partial}{\partial x_j} \left(\rho U_j C_p T \right) = \frac{\partial}{\partial x_j} \left(\alpha \frac{\partial T}{\partial x_j} - \rho C_p \overline{u_j T} \right) + S_i^T
$$
\n(3.4)

$$
\rho C_p \overline{u_j T} = -\frac{\mu_r}{\text{Pr}_i} \frac{\partial T}{\partial x_j}
$$
\n(3.5)

Where C_p is the specific heat capacity, α is the thermal conductivity, S_i^T is the energy per unit volume added to the gas phase by the influence of particles. Pr, is the turbulent Prandtl number and defined as 0.9.

3.2.2 Turbulence Model

Fluctuations in the transport quantities such as momentum, energy, and species are the characteristics labeled by the turbulence flow. These kinds of fluctuation can be small or large depending on the description of practical use in the simulation calculation. It is well known that no single turbulence model is universally acceptable for all classes of problems. So the selection of the turbulence model largely depends on consideration of the physies properties in a speeifie problem. In Fluent, there are several seleetions for the turbulence model, and the purpose of this section is to analyze the computational efforts, cost in terms of CPU time and memory, and also the appropriation of deseribing the flow on the selection of realizable $\kappa - \varepsilon$ turbulence model in this problem.

Different methods of ealeulating the turbulenee flow have been proposed in the literature review research in Chapter 1. And the $\kappa - \varepsilon$ turbulence model is presently the most widely used and validated of the two-equation models. It is determined from the solution of transport equations for the turbulenee kinetie energy and the energy dissipation rate, and produees a relatively satisfaetory result for the results of eomplex flows. This teehnique is available for the wall funetions to treat the near-wall sub-layers. The realizable $\kappa - \varepsilon$ model [28] is a relatively recent turbulence model. The term "realizable" means that the model is eonsistent with the physies of turbulenee flows and satisfies certain mathematical constraints. Neither the standard $\kappa - \varepsilon$ model nor the RNG $K - \varepsilon$ model is realizable. Compared with them, the realizable $K - \varepsilon$ model provides superior performanee for flows involving rotation, boundary layers under strong adverse pressure gradients, reeireulation, and separation. It also has the immediate benefit of predieting the spreading rate of both planar and round jets more aeeurately [37]. Finally, the performance of realizable $\kappa - \varepsilon$ model has been found to be substantially better than that of the other models.

In the realizable $\kappa - \varepsilon$ model, a new realizable eddy viscosity formulation and a new transport equation for the dissipation rate ε are derived from an exact equation. The eddy viscosity is no longer constant in the realizable $\kappa - \varepsilon$ model, and it relies on the mean strain and rotation rates, the turbulence fields, and the angular velocity of the system rotation. The modeled transport equations for κ and ε in the realizable $\kappa - \varepsilon$ model are expressed as follow:

$$
\frac{\partial}{\partial x_j} \left(\rho k U_j \right) = \frac{\partial}{\partial x_j} \left[\left(\mu + \frac{\mu_r}{\sigma_{\varepsilon}} \right) \frac{\partial k}{\partial x_j} \right] + G_k + G_b - \rho \varepsilon \tag{3.6}
$$

$$
\frac{\partial}{\partial x_j} \left(\rho \varepsilon U_j \right) = \frac{\partial}{\partial x_j} \left[\left(\mu + \frac{\mu_r}{\sigma_s} \right) \frac{\partial \varepsilon}{\partial x_j} \right] + \rho C_1 S_\varepsilon - \rho C_2 \frac{\varepsilon^2}{k + \sqrt{\nu \varepsilon}}
$$
(3.7)

In these equations, G_k represents the generation of turbulence kinetic energy due to the mean velocity gradients, $G_k = -\rho u_i u_i + G_h$ is the generation of turbulence **cbr.** kinetic energy due to buoyancy, calculated as: $G_h = -g_i \frac{\mu_i}{R} \frac{\partial P}{\partial r}$. Here β ρ Pr_{q} $C_1 = \max \left[0.43, \frac{\xi}{\xi + 5} \right], \xi = S \frac{k}{\epsilon}, S = \sqrt{2S_{ij}S_{ij}}$ is the modulus of the mean rate strain tensor, and $C_2 = 1.9$. The terms σ_k and σ_s are the turbulent Prandtl numbers for κ and ε , σ_k =1.0, σ_{ε} =1.2. The inlet boundary condition of *k* and ε depend on the turbulence intensity *I* , turbulence length scale *I* and the inlet velocity, expressed as:

$$
\kappa = \frac{3}{2} (u \cdot I)^2 \tag{3.8}
$$

$$
\varepsilon = \frac{C_{\mu}^{\frac{3}{4}}k^{\frac{3}{2}}}{l} \tag{3.9}
$$

Here $I = 0.16 \text{Re}^{-1/6}$ and $I = 0.07D_h$, Re is the Reynolds number based on the length of the top inlet. In this research configuration, the turbulent intensity is 5%. The hydraulic diameter D_h is 0.1 m.

3.2.3 Equations of Particles

As mentioned in Chapter 1, the numerical model uses a two-way coupled Euler-Lagrange method. The trajectories of the particles are predicted by integrating the force balance in the Lagrange reference frame. So the force balance which equals the particle inertia force with the force acting on the solid particle can be written as:

$$
\frac{du_{P,j}}{dt} = \frac{18\mu}{\rho_P d_P^2} \frac{C_D \text{Re}_{P,j}}{24} \left(u_i - u_{P,j} \right) + \frac{g_i (\rho_P - \rho)}{\rho_P} \tag{3.10}
$$

where $u_{p,i}$, ρ_p , and d_p are the velocity, density and diameter of the particle, respectively. $\text{Re}_{P,i}$, the relative Reynolds number, can be written as:

$$
\operatorname{Re}_{P,i} = \frac{\rho d_P |u_{P,i} - u|}{\mu} \tag{3.11}
$$

and C_p , the drag force coefficient, can be expressed as:

$$
C_D = \frac{24}{\text{Re}_{p_i}} \left(1 + 0.15 \,\text{Re}_{p_i}^{2/3} \right) \tag{3.12}
$$

In this case, Re_{p_i} is up to 10³, so the drag coefficient can be expressed above by the method introduced by Clift et al. [38]. Because the material and gas density ratio is very large, the effect of the pressure gradient and particle acceleration on drag coefficient can be neglected. Here, the particle volume fraction is very small and the diameter of the particle is much smaller than the distance between particles, so the particle-particle drag influence and collision can be neglected during the calculation.

The energy equation for the particle is given as:

$$
\frac{\rho_p C_p \pi d_p^3}{6} \frac{dT_p}{dt} = Nu \alpha_p \pi d_p (T - T_p) + \pi d^2 \varepsilon_p \sigma (T_n^4 - T_p^4)
$$
\n(3.13)

where the *Nu* is the Nusselt number defined by

$$
Nu = 2 + 0.6 \,\text{Re}_{p}^{1/2} \,\text{Pr}^{1/3} \tag{3.14}
$$

here Pr is the Prandtl number of the gas flow, ε_p is the particle emissivity, σ is the Stefan-Boltzman constant, α_p is the thermal constant of the particle, and T_R is the radiation temperature, which is determined by :

$$
T_R = (G/4\sigma)^{1/4} \tag{3.15}
$$

here *G* is the incident radiation value,

$$
G = \int_{\Omega = 4\pi} d\Omega \tag{3.16}
$$

3.2.4 Radiation Model

The radiation model used in this study considered the interaction between the radiation field and the falling particle curtain. To keep the accuracy of the calculation, a discrete ordinate (DO) method with a solar ray tracing model is used inside the solar receiver. The DO radiation model solves the radiative transfer equation for a finite number of discrete solid angles, each of them associated with a direction noted by *s* in the global Cartesian coordinate. The DO model solves for as many transport equations as there are directions \vec{s} . The radiative transfer equation can be written as:

$$
\nabla \cdot (I_{\lambda}(\vec{r}, \vec{s})\vec{s}) + (a_{\lambda} + \sigma_{s})I(\vec{r}, \vec{s}) = a_{\lambda}n^{2}I_{b\lambda} + \frac{\sigma_{s}}{4\pi} \int_{0}^{4\pi} I_{\lambda}(\vec{r}, \vec{s}')\Phi(\vec{s} \cdot \vec{s}')d\Omega' \tag{3.17}
$$

 $I_{\lambda}(\vec{r},\vec{s})$ is the spectral intensity in direction \vec{s} and position \vec{r} of an absorbing, emitting, and scattering medium. λ is the wavelength, a_{λ} is the spectral absorption coefficient, σ_{s} is the scattering coefficient, \vec{s}' is the scattering direction vector, Φ is phase function, Ω' represents solid angle, and $I_{b\lambda}$ is the black body intensity given by the Planck

function. The total intensity is computed as $I(\vec{r},\vec{s}) = \sum I_{ik}(\vec{r},\vec{s})\Delta\lambda_k$, where the *k* summation is over the wavelength bands.

The solar ray tracing algorithm can be used to predict the direct illumination source which results from the direct solar incidence. It takes a beam that modeled the position of the sun vector and illumination parameters, applies it to any wall or inlet/outlet boundary conditions, and performs a face-by-face shading analysis to determine well-defined shadows on all boundaries.

3.2.5 User Defined Function on the Drag Force

In the discrete phase model, particles interact with a continuous phase. A user defined function on the drag force is applied on each particle to evaluate the particle behavior under the influence of heat buoyancy force and gravity force balance.

3.3 Numerical Modeling Approach

A three-dimensional model of the solid particle solar receiver (SPSR) with the influence of an air-jet is analyzed. The free-falling down particles will form a solid particle curtain and be directly heated by the reflected solar energy that passes through the aperture of the cavity. The mass, momentum and energy exchange between the solid particle phase and gas fluid phase are simulated by the two-way coupling Euler-Lagrange method. A discrete ordinate radiative transfer method has been applied to study the coupling of radiative heat transfer and the falling particle curtain. The realizable $\kappa - \varepsilon$ model is used in the investigation of turbulence flow. In order to predict the performance of the SPSR, the aerodynamic behavior of the particles and thermal interaction, which include particle-particle radiation, particle-wall radiation, particle-air convection, and air-wall convection are analyzed and demonstrated in this work. The numerical results of the parametric studies will be discussed in Chapter 4 and Chapter 5.

3.4 Solution Algorithm

3.4.1 Pressure-Velocity Coupling: SIMPLE

FLUENT provides several kinds of pressure-velocity coupled methods. For this problem, increasing the pressure-correction under-relaxation to 1.0 can lead to instability, because of the huge grid size. Under this situation, a slightly conservative underrelaxation is needed, and a selection of SIMPLE algorithm is necessary. The acronym SIMPLE stands for Semi-lmplicit Method for Pressure-Linked Equations. This algorithm is essentially a guess-and-correct procedure for the calculation of pressure on the staggered grid [23].

3.4.2 Second-Order Upwind Scheme

A volume discretization scheme stores discrete values of the scalar and values at faces must be interpolated from the cell center value. This is accomplished by using an upwind scheme. Upwind means that the value at interface is derived from the data in upstream cells. There are several upwind schemes, and a second-order upwind scheme is applied in this problem. In this scheme quantities at the cell face are computed using a multidimensional liner reconstruction approach [39]. Higher-order accuracy is achieved at the cell interface through the Taylor series expansion of the cell-centered solution. Second-order upwind is available in the pressure-based and density-based solvers.

3.4.3 PRESTO Scheme

The PRESTO (pressure staggering option) scheme is the discretization scheme for the pressure in flows with high swirl number, high-Rayleigh-number natural convection, and high-speed rotating.

3.4.4 Under-relaxation Factors

Under-relaxation factors contain the factors for all equations in the pressure-based solver. It includes turbulence variables and viscosities, which are solved separately in all equations. In the beginning, the under-relaxation parameters are set to be default values for calculation. In this complicated natural convection problem the density is strongly coupled with temperature, so the under-relaxation factors can be adjusted under default values to achieve convergent results.

3.5 Boundary Conditions

Before the solid particles fall down from the top opening of the cavity, along each trajectory, the particles have an initial temperature of 873 K, and are given an initial falling speed of 0.088 m/s. A two-band isotropic radiation calculation is made with a solar band from 0.2 to 1.5 micron and an infrared band from 1.5 to 30 micron. An incident solar flux of 920 suns is applied through the open aperture of the cavity.

3.5.1 Pressure Outlet Boundary Condition

Pressure outlet boundary conditions require the specification of a static pressure at the outlet boundary. The boundary containing the area around the SPSR is set as pressure outlet boundary. A set of "backflow" conditions are specified, because reverse flow should occur at the exit during the calculation. At the pressure outlet, a constant zero gauge pressure is entered, with a total backflow temperature of 300 K. In order to

determine the direction of the backflow, the direction of the flow in the cell layer adjacent to the pressure outlet is selected.

3.5.2 Velocity Inlet Boundary Condition

As mentioned earlier, the air-jet has the potential of decreasing the convection and radiation loss from the SPSR. So the velocity inlet boundary condition is applied in the air-jet, which is on top of the receiver aperture. The velocity inlet boundary condition is used to define the flow velocity, along with all other relevant scalar properties of the flow. This kind of boundary is intended to be used in incompressible flow. In this research, a velocity specification method defined the magnitude and the direction of the air-jet.

3.5.3 Wall Boundary Condition

Wall boundary conditions are used to bound gas and particle flow in this study. Since the flow is modeled as viscous flow, the no-slip boundary condition is enforced at the wall. The heat flux is assumed to be zero (i.e. adiabatic) in the present research. The wall of the cavity with an internal emissivity of 0.8 and absorptivity of 0.9 are used for both bands (visible and infrared). The bottom wall can trap the falling particles, while the side walls reflect the particles striking on them.

3.6 Physical Properties Used in the Numerical Model

The detailed information related to the physical prosperities of the solid particles and gas flows in the numerical model are listed in Table 3.1.

Gas Flows	Solid Particle
Inlet Air Temperature of air-jet: 300-700 K	Particle Density: 3200 kg/m^3
Operating Pressure: 101000 Pa	Heat Capacity: 1085 J/kg-K
Incompressible Ideal Gas	Thermal Conductivity: 6.67 W/m-K
	Particle Diameter: 60-800 micron

Table 3.1 Operating conditions of gas flow and physical properties of the solid particle

3.7 Mesh Independent Study

In this research, GAMBIT is used as a pre-processor to create a computational mesh. Several computational meshes have been generated to check the grid independence. Meshes with 197,892, 332,960, 533,400, 717,094, and 909,441 cells have been investigated. Figure 3.1 gives the positions of the lines and points, along which data are plotted and compared. Figure 3.2 shows the velocity distribution along line 1 ($x=1$ m, $y=1.4$ m, $z=0$ m) to $(x=1$ m, $y=1.4$ m, $z=1.58$ m) and Figure 3.3 shows the velocity distribution along line 2 (x=0 m, y=0.7 m, Z=0.8 m) to (x=2 m, y=0.7 m, z=0.8 m). Both figures present very uniform velocity distribution tendency and magnitudes, except for the mesh with smallest cell number.

Figure 3.1 The positions of line 1 and line 2, point A and point B

Figure 3.4 and Figure 3.5 show the velocity relative error percentages at point A $(x=1 \text{ m}, y=1.4 \text{ m}, z=0.3 \text{ m})$ and point B $(x=0.15 \text{ m}, y=0.7 \text{ m}, z=0.8 \text{ m})$. The results conclude that the mesh with 533,400 cells, as shown in Figure 3.6, has velocity relative error percentages under 10 % for both the point value, compared with the finest mesh. To save the computational resources and time, a computational mesh with 533,400 cells **is** used **in the continuing research work.**

Refined and gradient mesh is designed in the boundary, open aperture and particle curtain region of the SPSR, as shown in Figure 3.7. This can provide an accurate result for the place that has larger pressure, temperature and velocity differences.

Figure 3.2 Velocity magnitude along line 1

Figure 3.3 Velocity magnitude along line 2

Figure 3.4 Velocity relative error compares to the finest mesh at point A

Figure 3.5 Velocity relative error compares to the finest mesh at point B

Figure 3.6 Computational mesh of the SPSR

CHAPTER 4

NUMERICAL SIMULATION OF A THREE-DIMENSIONAL SOLID PARTICLE SOLAR RECEIVER

Numerical modeling of the SPSR was done by SNL more than 20 years ago. All the parametric studies which are based on the two-dimensional design do not consider the influence of an air-jet. Three-dimensional computational fluid dynamics studies of the SPSR were first studied by Chen et al. [30], to compare cavity efficiency and average exit particle temperature with and without bottom opening. And the influence of air-jet on SPSR was studied by Tan et al. [40] and Chen et al. [41].The effect of an air-jet has been studied and proved to reduce the convection loss of a central solar receiver, by isolating the interior from the surrounding environment [9]. In this solid particle solar receiver study, the air-jet has the additional benefit of protecting the spreading of the solid particles, which are light and could leak to the outside with heat energy. In a practical design, the SPSR will be set up on a very high solar tower, where the wind from different directions in that elevation can blow the heat media solid particles out of the receiver, if there is no protection. The characteristics of the air-jet based on the conceptual design can be investigated by performing parametric studies related to the cavity efficiency and the exit particle temperature. The results can provide the fundamental engineering information for the solid particle solar receiver scale-up design.

4.1 Comparison of SPSR with and without an Air-jet

Typical air flow patterns with and without an air-jet in the selected slice $(x=1 \text{ m},$ center plane of the receiver) are shown in Figure 4.1. Here, particle diameter is chosen as 650 micron, particle initial temperature is given as 873 K, particle mass flow rate is 5 kg/s, and particle initial downward speed is 0.088 m/s. A uniform solar incident radiation flux of 920 suns is considered in front of the solid particle curtain. Both the particles and the back wall are heated by the solar radiation. As a result, a buoyancy force is generated around the particle curtain. There is a strong air circulating occurring in the region between the front aperture and particle curtain, as shown in Figure 4.1 (a). A large amount of air is entering into the cavity through the bottom of the aperture from outside. As the temperature of the air entering into the receiver through the aperture is much lower than that inside the receiver, the cavity efficiency and exit particle temperature will be reduced due to the energy loss by convection.

(a) Without an air-jet

(b) Air-jet velocity is 8 m/s

Figure 4.1 Air flow pattern without an air-jet (a) and with an air-jet (b)

Figure 4.1 (b) illustrates the effect of applying an air-jet to cover the open aperture. There still **exists** a strong clockwise circulation between the aperture and particle curtain. But the injected downwards air-jet wind blocks this circulation from moving outwards. Small air exchanges between the cavity and outside field can happen by convection. This is the expected air flow pattern for SPSR designers. Figure 4.2 shows the temperature (K) contours at slice $x = 1$ m (in the centerline of the receiver). Both designs can reach the highest temperature of 2100 K at the back walls by absorbing the incident solar flux. But temperature inside the receiver is reduced in Figure 4.2 (a) by sucking cold air flow from outside. For the case with protection from an air-jet, as shown in Figure 4.2 (b), the air temperature inside cavity can keep a relatively high value.

(b) Air-jet velocity is 8 m/s

The thermal performanees of the SPSR are characterized by the cavity efficiency and average exit particle temperature. The cavity efficiency is introduced as:

$$
\eta = \frac{Q_P}{Q_{inc}}\tag{4.1}
$$

Where Q_p is the absorption energy gain of the solid particles, and Q_{inc} is the total incident solar energy that comes through the aperture. Obviously, the total incident solar energy value is constant if the solar incident flux doesn't change. The radiation loss percentage is defined as the ratio of radiation heat loss $Q_{radloss}$ and total incident solar energy Q_{inc} , the same as convection heat loss percentage, which is defined as the ratio of convection heat loss Q_{envloss} and total incident solar energy Q_{inc} .

Table 4.1 compares the cavity efficiencies, average exit particle temperatures, heat convection and radiation loss percentages with and without air-jet. It clearly shows the performance improvement of SPSR after applying an air-jet.

4.2 Parametric Studies of SPSR

In section 4.1, comparisons of the SPSR with and without air-jet are made to evaluate the benefit of an air-jet. But the designs of an air-jet, like air-jet velocity and airjet temperature, have not been considered. In this part, several parametric studies are performed to identify the optimized operating conditions for SPSR.

4.2.1 Aerodynamic Analysis of SPSR with Different Air-jet Velocities

The SPSR with an air-jet of 8 m/s performs better than the one without an air-jet, as shown in section 4.1. A range of 0 m/s $-$ 10 m/s air-jet velocities are numerically tested to find the optimized velocity result in this numerical model, with all the same working conditions of SPSR except air-jet velocity. Mass flow rate is 5 kg/s, downward air-jet temperature is 300 K, mass flow rate is 5 kg/s, and solar flux value is 920 suns. From Figure 4.3 to Figure 4.6, including Figure 4.1 and Figure 4.2, all the air flow patterns and temperature contours at the central plane of the receiver give a very reasonable explanation of the result that is concluded in Figure 4.7. In Figure 4.7, the cavity efficiencies and average exit particle temperatures keep on increasing from air-jet velocity of 0 m/s upto 8 m/s. For the cases of air-jet velocity magnitudes less than 8 m/s, the air flow patterns show that air-jets are not powerful enough to cover the whole aperture. But as the air-jet velocity increases, much more heated air is protected inside the cavity. The temperature contours also show the air encirclement tendency inside the cavity. From Figure 4.7, air-jet velocity of 8 m/s provides a good screening result, better than 10 m/s. In this investigation of optimal working velocity, each air-jet has an initial temperature of 300 K, which is lower than the air temperature inside the cavity. When velocity of an air-jet reaches a high value, it might blow the cold wind of the air-jet into

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the cavity. This explains that selection of 10 m/s air-jet is not necessary in this numerical model, even through it covers the whole aperture as shown in Figure 4.6.

(a) Air flow pattern

(a) Air flow pattern

(a) Air flow pattern

Figure 4.7 Cavity efficiencies of SPSR and average exit particle temperatures as a function of different air-jet velocities

Figure 4.8 shows the radiation loss and convection loss percentages as a function of different air-jet velocities. The radiation loss percentage value is almost constant in the tested air-jet velocity range of $0 \text{ m/s} - 10 \text{ m/s}$, which occupies about 30 % of the total incident solar energy. The convection loss percentage keeps on decreasing until the air-jet velocity reaches 8 m/s, which is the smallest value of convection loss and also gives the best perfomiance of SPSR in the tested velocity range. These phenomena prove that an air-jet is capable of reducing convection heat loss.

Figure 4.8 Radiation loss and convection loss percentages as a function of different air-jet velocities

4.2.2 The Influence of Different Air-jet Temperatures on SPSR

Since a temperature of 300 K is a low initial value for an air-jet, it may inject this cold air into the SPSR. An assumption of increasing initial temperature of an air-jet to obtain a better performance of the SPSR is made. To save the energy cost for running the air-jet, the pre-heat air, which exists around high temperature particles, can be recuperated. Through this method, the initial temperature of the air-jet can reach about 700 K. Other working conditions are the same: the downward air-jet velocity is 8 m/s, particle diameter is 650 micron, mass flow rate is 5 kg/s, and solar flux value is 920 suns. negTo characterize the air flow clearly, the air flow stream lines, which are released from the air-jet, are shown in Figure 4.9. In this case the air-jet velocity is 8 m/s, and the initial temperature is 700 K. The stream lines with high temperature are pushed forwards, and cause a leakage in the bottom of the aperture. Cold air from surroundings is drawn into the cavity of the SPSR. This figure also shows air flow pattern and temperature contours in the centerline of the receiver, which demonstrate the backflow of cold air clearly. The air-jet stream lines in Figure 4.10, which has the same air-jet velocity but with an initial temperature of 300 K, shows that the air curtain injected by the air-jet is able to seal the open aperture. As a result, the increased temperature of the air-jet can not reach a better performance than the one of low temperature. The pre-heat air from the air-jet causes a buoyancy force, which forms a strong backflow of air and pushes the air stream lines from the air-jet away from the aperture. If the temperature of the air-jet is not as high as 700 K, the buoyancy force may not be that strong. The result of air-jet with lower temperature is shown in Figure 4.11, but there still exists leakage. The air-jet velocity of 8 m/s may be not strong enough to balance that buoyancy force. Figure 4.12 shows steam lines of a stronger air-jet flow, where the velocity is 10 m/s and the temperature is 700K. No great improvement of the performance has been found through this design.

Figure 4.9 Air flow stream lines released from air-jet. The stream line is colored by temperature. Air-jet velocity is 8 m/s and air-jet temperature is 700 K.

Figure 4.10 Air flow stream lines released from air-jet. The stream line is colored by temperature. Air-jet velocity is 8 m/s and air-jet temperature is 300 K.

Figure 4.11 Air flow stream lines released from air-jet. The stream line is colored by temperature. Air-jet velocity is 8 m/s and air-jet temperature is 500 K.

Figure 4.12 Air flow stream lines released from air-jet. The stream line is colored by temperature. Air-jet velocity is 10 m/s and air-jet temperature is 700 K.

Table 4.2 summarizes the performance of the SPSR with air-jet temperature in the range of 300 K to 700 K, and a velocity range of 6 m/s to 10 m/s. All results don't show big differences, so the different temperatures application of a downward air-jet won't influence the performance of the SPSR.

Velocity of Air-jet (m/s)	6		8		10	
Temperature of Air-jet (K)	η	T_{P-exit} (K)	η	T_{P-exit} (K)	η	T_{P-exit} (K)
300	0.61	1107	0.62	1108	0.61	1106
500	0.61	1105	0.60	1104	0.62	1111
700	0.60	1100	0.61	1105	0.61	1106

Table 4.2 Cavity efficiencies and average exit particle temperatures with different air-jet velocities and different air-jet temperatures

4.2.3 Air-jet Injected from the Bottom of Open Aperture

A model of applying a downward pre-heated air-jet to cover the open aperture is proved to be useless on improving the performance of the SPSR. A buoyancy force is generated by temperature differences between the air-jet and surrounding area, which pushes the air-jet forward and causes leakage. An air-jet injected upward from the bottom of the aperture is investigated, such that a buoyancy force can act in the same direction as the air flow of the air-jet. From the air flow pattern of the air-jet without being heated, as shown in Figure 4.13 (a), the air from the air-jet blows into the cavity, distorts the flow to lower the cavity efficiency. From Figure 4.13 (b), the temperature contour describes the 400 K isotherm inside cavity, which well demonstrates the invasion of cold air. Under this situation, the cavity efficiency is $\eta = 60\%$, and average exit particle temperature is $T_{p,ew} = 1102K$. This case showed the air flow behavior without considering the heat buoyancy force.

The same model with initial air-jet temperature of 700 K is studied, and this temperature of air-jet is much higher than the surrounding temperature of 300 K. As a result, a floating force is created to hold the injection direction of air-jet. From Figure 4.14 (a), the air flow pattern illustrates the well formed air-jet curtain, which protects the heat inside the cavity. Even air from air-jet injects into the cavity, which is shown from the 700 K isotherm in Figure 4.14 (b), this high value of temperature cannot influence the performance inside the cavity. Here, cavity efficiency increases to $\eta = 62\%$, and average exit particle temperature is 1110 K. From Figure 4.13 (c) and Figure 4.14 (c), the particle distributions on the bottom plane appear to be steady and uniform. Compared to the performance of air-jet injected from top, preheat air-jet injected from bottom of open aperture doesn't improve much on the performance of SPSR.

(b) Temperature contour at slice x=l m

(c) Particle distributions on the bottom plane, colored by temperature (K)

Figure 4.13 Air-jet injected from bottom of aperture. The velocity of air-jet is 8 m/s, temperature is 300 K, mass flow rate is 5 kg/s, diameter of partiele is 650 micron, and solar flux is 920 suns

(a) Air flow pattern at slice $x=1$ m

(c) Particle distributions on the bottom plane, colored by temperature (K)

4.2.4 The Influence of Different Particle Diameters on SPSR

Chen et al. [30] show the heated particle curtain with smaller diameter does not remain as flat as the larger ones, and is more capable of reaching higher temperature. Hruby's [32] research results also conclude that smaller particle size can bring both higher efficiency and exit temperature. The present study considers the influence of an air-jet. Cavity efficiency and average exit particle temperature relate to the eharaeteristies of the solid particles. The trajectory of the solid partiele is predicted by integrating the force balance on the partiele. The force balance on each particle equals the particle inertia force with the force acting on each particle. Buoyancy force acts more predominantly on particles with smaller diameter than larger ones, which helps smaller particles to stay in the cavity for a longer time, and accept more solar irradiation. Based on the model of a two-way coupled Euler-Lagrange method, the calculation includes the exchange of heat and momentum between the gas phase and particles. But it neglects the effect of the nearby particles on the drag force. This assumption is reasonable when the partiele volume fraction is very small and the particle distance is much larger than the diameter of the particles. From Figure 4.15 and Figure 4.16, the partiele volume fraction at the center slice $(x=1m)$ of the cavity shows that the particle fractions are small enough to neglect the particle-particle force at different heights.

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Figure 4.15 Particle volume fraction as a function of distance from back wall at different height in select slice $(x = 1$ m). The diameter of the particles are 200 micron, and air-jet velocity is 8m/s. H is the distance from the bottom wall.

Figure 4.16 Particle volume fraction as a function of distance from back wall at different height in select slice $(x = 1$ m). The diameter of the particles are 650 micron, and air-jet velocity is 8m/s. H is the distance from the bottom wall.

Figure 4.17 gives the result of the SPSR cavity efficiency and average exit particle temperature as a function of different particle diameters with an air-jet velocity of 8m/s. Applying a smaller particle size in the SPSR shows a higher average exit particle temperature, resulting in higher cavity efficiencies. The smaller particles might spread due to strong turbulence air flow pattern inside the cavity, which will cause particle and energy loss. From the particle volume fraction figures, this possibility can be neglected. Most of the particles stay to form a relatively uniform curtain without any particle mass loss, until the diameter of particles decreases to 60 microns, which is too small to stay inside the cavity, and particles start to leak from the front aperture. So theoretically, the partiele with diameter in the range of 70-80 micron gives the best performance of SPSR. And the cavity effieieney begins to decrease when the diameter is less than 70 micron, because of the heat energy carried away by the leaking particles. On the other hand, particles that keep inside cavity still increase their average exit temperature, because of the longer residence time.

Figure 4.18 and Figure 4.19 show the particle distributions on the bottom plane in the eases of partiele diameter 200 micron and 650 micron, colored by the temperature. 700 particles are tracked on each calculation step, so the particles with different size appear the same number on the bottom. Over time, the solid partiele diameter of 200 micron case will have more particles falling to the bottom. The partiele distributions on the bottom show the curtains have not been pushing towards the front or back of the cavity, which means the air-jet is very effective at protecting the particle flow inside cavity. From the particle distribution colored by temperature, the smaller partiele size of 200 micron ease reaches higher average exit temperature.

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Figure 4.17 Cavity efficiency and average exit particle temperature as a function of particle diameters, velocity of air-jet is 8 m/s, downward air-jet temperature is 300 K, mass flow rate is 5 kg/s, and solar flux is 920 suns

Previous analysis concludes that smaller particle size can bring higher cavity efficiency and average exit particle temperature. Theoretically, the optimized size of the solid particle is 70-80 micron, and particles with diameter less than this range will spread out of the cavity. This proves the existence of a limitation on the size of the particle, because small particles will spread, and distort the particle curtain. The simulation analysis provides the optimized SPSR working condition on the particle size, but this ideal small particle size might be unpractical for manufacture, and cause agglomeration and fracture under extremely high temperature, which largely depends on the characteristics of the particle material.

Figure 4.18 Particle distributions on the bottom plane. Particle size is 200 micron, and air-jet velocity is 8 m/s.

Figure 4. 19 Particle distributions on the bottom plane. Particle size is 650 micron, and air-jet velocity is 8 m/s.

Figure 4.20 is the particle distributions on the bottom plane, colored by temperature with the particle diameter of 100 micron. Compared to the case of 200 micron and 650 micron, the distribution pattern turns out to be a curve that is pushed to the back of the wall. The track of the particle on the bottom is distributed more widely, and with a relatively uniform temperature. Several particles are attracted to the wall. This phenomenon has been found by using 100 micron diameter of the particle. It is because the back flow drag-force and buoyancy force hold the particle inside the cavity for a longer time and some particles cannot hold the starting falling projection position when they fall to the bottom.

Figure 4. 20 Particle distributions on the bottom plane. Particle diameter is 100 micron, and air-jet velocity is 8 m/s

Higher cavity efficiency and average exit particle temperature of the SPSR is achievable by decreasing the diameter of solid particles. While the size of particle is too small, it is unable to stay inside the cavity under the influence of strong buoyancy force. And leakage of high temperature particles will cause thermal energy loss. So it is necessary to find out the optimized range of partiele size, which not only brings the highest average exit particle temperature but also keeps all the particles inside the cavity. Figure 4.21 and Figure 4.22 show the particle distribution on the bottom plane when the diameter of the solid particles deceases to 80 and 70 micron. And other working conditions of the SPSR are the same as previous cases with larger size of solid particles. The downward air-jet velocity is 8 m/s, air-jet temperature is 300 K, mass flow rate is 5 kg/s, and solar flux value is 920 suns. At these conditions, the particles still stay inside the cavity with a better performance compared to larger size of the solid particles. Figure 4.23 is the particle distribution on the bottom plane with a diameter of 60 micron, and it shows the spreading of particles. It is been told that the cavity efficiency starts to decrease when the diameter of solid particles is less than 70 micron, because the heated particles start to leak out of the cavity from the aperture, and this causes heat energy loss. On the other hand, the average exit particle temperature of SPSR with partiele diameter of 60 micron is still increasing. This result just calculates the temperature of the solid particles that are trapped on the bottom plane and does not consider the other leaking particles. Each trapped small particle can absorb more heat energy, compared to larger size of solid particles.

Figure 4. 21 Partiele distributions on the bottom plane. Particle diameter is 80 micron, and air-jet velocity is 8 m/s

Figure 4. 22 Particle distributions on the bottom plane. Partiele diameter is 70 micron, and air-jet velocity is 8 m/s

Figure 4. 23 Partiele distributions on the bottom plane. Partiele diameter is 60 micron, and air-jet veloeity is 8 m/s

4.2.5 Aerodynamic Analysis of SPSR with Different Particle Mass Flow Rates

The eontrol of mass flow rates ean he realized by adjusting the opening size of the particle inlet. And the volume percentage of partiele inside the eavity is still under 1 % assumption when the mass flow rate is 7 kg/s. It promises there is no particle-partiele eollision in this mass flow rate ealeulation range. So the simulation model is aeceptahle. Figure 4.24 demonstrates that inereasing partiele mass flow rate leads to inereased eavity effieieney and deereased average exit partiele temperature. Larger mass flow rates can bring more partieles to transport absorbed solar energy, whieh will inerease the overall efficiency. On the other hand, more particles will increase the thickness and reduce the transparency of particle curtain, which prevents some particles from absorbing the radiated and reflected energy, resulting in a deereased average exit partiele temperature.

Figure 4. 24 Cavity efficiency and average exit particle temperature as a function of different particle mass flow rates, downward air-jet temperature is 300 K, mass flow rate is 5 kg/s, diameter of particles is 650 micron, and solar flux is 920 suns

4.2.6 The Influence of Different Solar Irradiation Values on SPSR

The solar irradiation intensity is another important input to the simulation. It will be varied depending on different latitudes, longitudes, and elevations around the earth. Even in the same location, different season and time during each day will get different results. The hase modeling case uses the solar irradiation value of 920 suns. The parametric study focused on solar flux will provide important information for the time and location operation instructions on the SPSR. Figure 4.25 shows the changing tendency of the efficiency and exit temperature with 650 suns, 920 suns, and 1280 suns solar fluxes with the influence of air-jet velocities 6 m/s and 8 m/s. The intensified irradiation will obviously bring higher cavity efficiency and average exit particle temperature. And in an intensified irradiation working condition, the SPSR with an air-jet velocity of 6 m/s performs more effectively than air-jet velocity of 8 m/s, which means that the heat convection loss is higher in SPSR with an air-jet velocity of 8 m/s. This is because stronger air-jet winds blow in cold air.

Figure 4. 25 Cavity efficiency and average exit particle temperature as a function of different solar irradiation values, downward air-jet temperature is 300 K, diameter of particle is 650 micron, and mass flow rate is 5 kg/s

CHAPTER 5

CONCLUSIONS AND RECOMMENDATION

5.1 Conclusions

This research presents the following results on the simulation of a SPSR with the influence of an air-jet. Using a CED technique a parametric study of the aerodynamic behavior is investigated in this research. The simulation results are aimed at providing the useful information to lead the design of a high performance SPSR. This paper explains the following major problems:

- (1) Generate a three-dimensional numerical model for solid particle solar receiver.
- (2) Compare SPSR cavity efficiency and average exit particle temperature with and without an air-jet.
- (3) Identify the optimized operation conditions of the SPSR, including: air-jet velocity, air-jet temperature, particle size, particle mass flow rate, and solar flux value.

The numerical results obtained in this research provide a better idea for the design process and the experimental research. They also provide insight in to how to improve the design of the SPSR.

The computational fluid dynamics analysis of the SPSR without an air-jet and with different air-jet velocities is presented in this study. A two way coupling EulerLagrange method was used to simulate the gas phase and solid particle flow inside the cavity. The solar ray tracing and discrete ordinate model evaluate the flow pattern, velocity, and efficiency inside the SPSR. The profiles for demonstrating the gas and particle aerodynamics and thermal behaviors are studied in detail. The parametric studies on the air-jet influence on SPSR provide better ideas on the method to increase the cavity efficiency, by modifying the design to optimize the SPSR performance.

The efficiency of the particle receiver and the particle exit temperature will increase by applying an air-jet design. The application of an air-jet can block the heat convection loss compared to the case without an air-jet, for the purpose of reducing the convection and radiation loss. A velocity of 8 m/s for the air-jet provides fairly good screening result, over the selected range of 0 m/s to 10 m/s. Increasing the particle mass flow rate leads to increased cavity efficiency, and decreased average exiting particle temperature. Tlxrough applying smaller particle size in the SPSR, a higher exiting particle temperature and efficiency is achievable. And particle with diameter of 70-80 micron is the optimized size that gives the best performance of SPSR. This particle size with a downward air-jet temperature of 300 K, air-jet velocity of 8 m/s, mass flow rate of 5 kg/s, and under the solar radiation of 920 suns, can bring the cavity efficiency of 85 % and average exit particle temperature of 1199 K, which proves this conceptual design of SPSR can achieve the minimum temperature requirement for the thermo-chemical reaction. On the other hand, using a heated air-jet to cover the aperture cannot help to improve cavity efficiency. This phenomenon was caused by the buoyancy force which acted on the pre-heated air-jet. Also the pre-heat air-jet injected from the bottom of the cavity doesn't influence the performance of SPSR. Higher solar flux radiation will give a better performance on the heat absorption.

5.2 Future Work

For the future studies related to this research, a few improvements can be realized, and the simulation results can be more reasonable. The recommended improvements of SPSR for the future study are as follows:

(1) Consider the wind influence combined with the performance of an air-jet.

(2) The boundary condition on the cavity wall is more reasonable if one considers the conduction effect.

(3) Heat inside the cavity can be recuperated.

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- Taide Tan, Yitung Chen, and Zhuoqi Chen, "Wind Effect on the Performance of Solid Particle SOLAR Receivers: Part II-The Performance with an Aerowindow Protection," submitted to *solar energy,* June 2008, (under reviewing).
- Taide Tan, Yitung Chen, and Zhuoqi Chen, "Wind Effect on the Performance of Solid Particle SOLAR Receivers: Part I- The Models and the Performance without an Aerowindow Protection," submitted to *solar energy,* June 2008, (under reviewing).

Thesis Title:

Numerical Analysis of the Solid Particle Receiver with the Influence of an Air-jet

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